

The Enriched Xenon Observatory for double beta decay

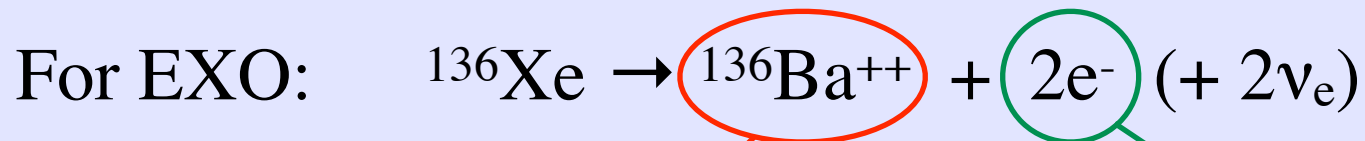
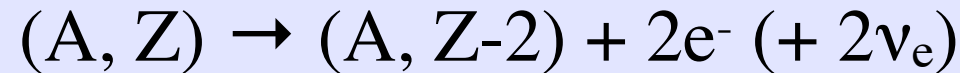
Andrea Pocar
Stanford University

Overview

- Neutrinoless double beta decay
- EXO concept
- The EXO-200 detector
- Barium ion identification

Double beta decay and EXO

Extremely rare decay of some nuclei:



positively identify daughter via
optical spectroscopy of Ba^+

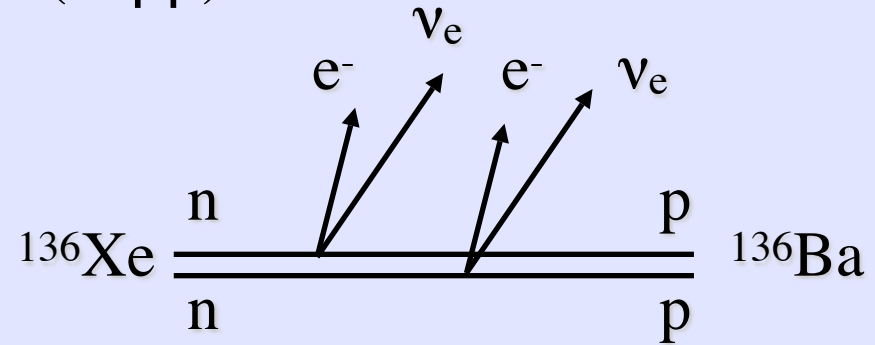
detect the 2 electrons
(ionization + scintillation
in xenon detector)

[M. Moe, Phys. Rev. C 44 (1991) R931]

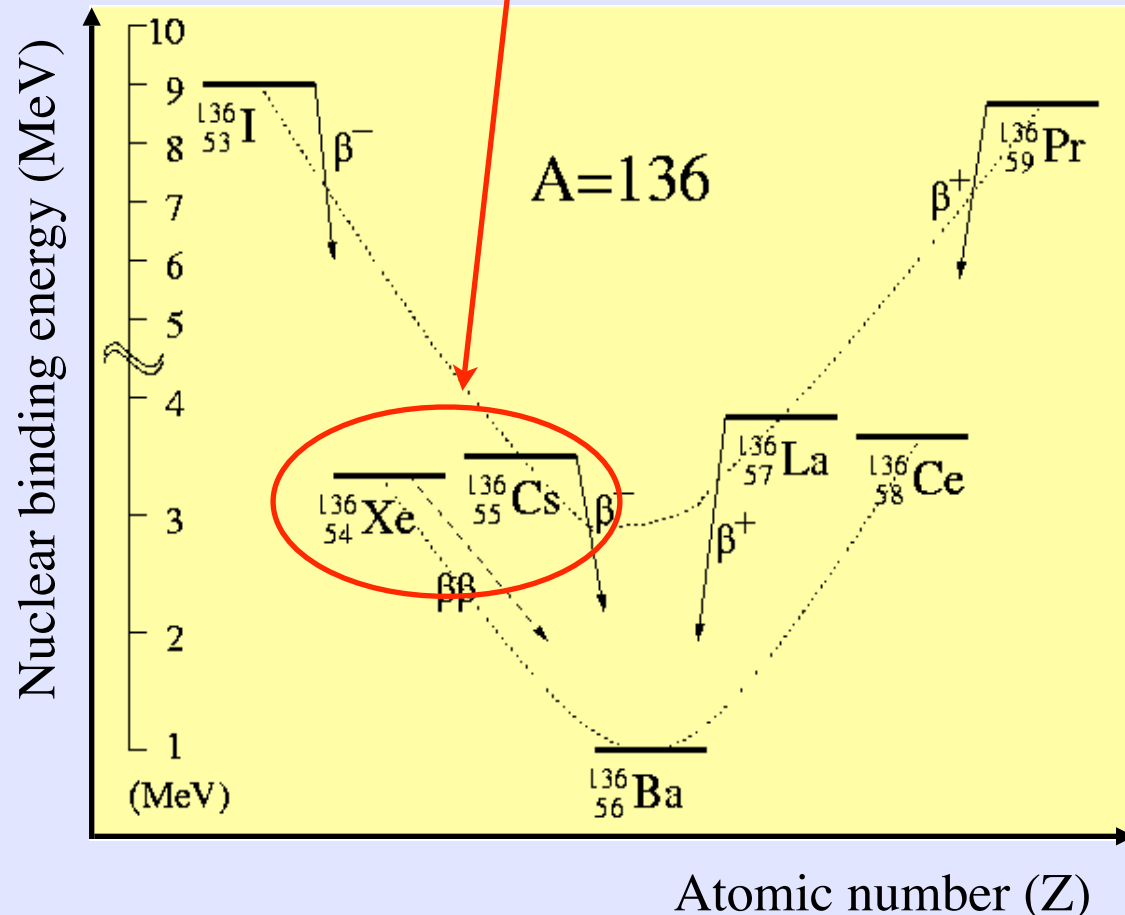
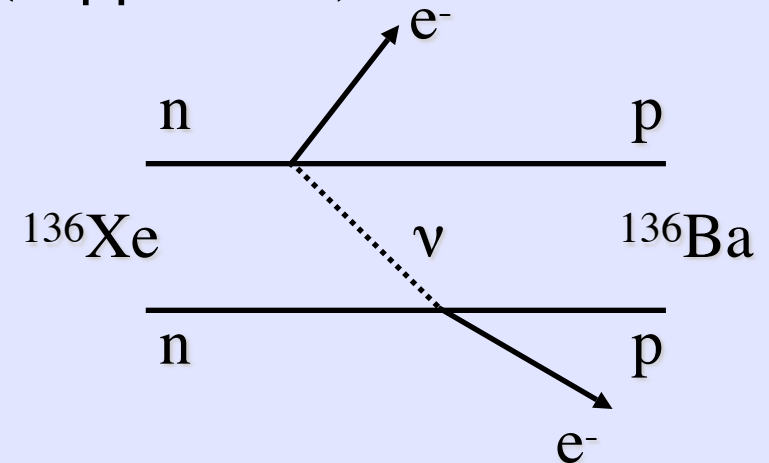
Double beta decay

a second-order process detectable when single β -decay is energetically forbidden (even-even nuclei)

standard electroweak process ($2\nu\beta\beta$):

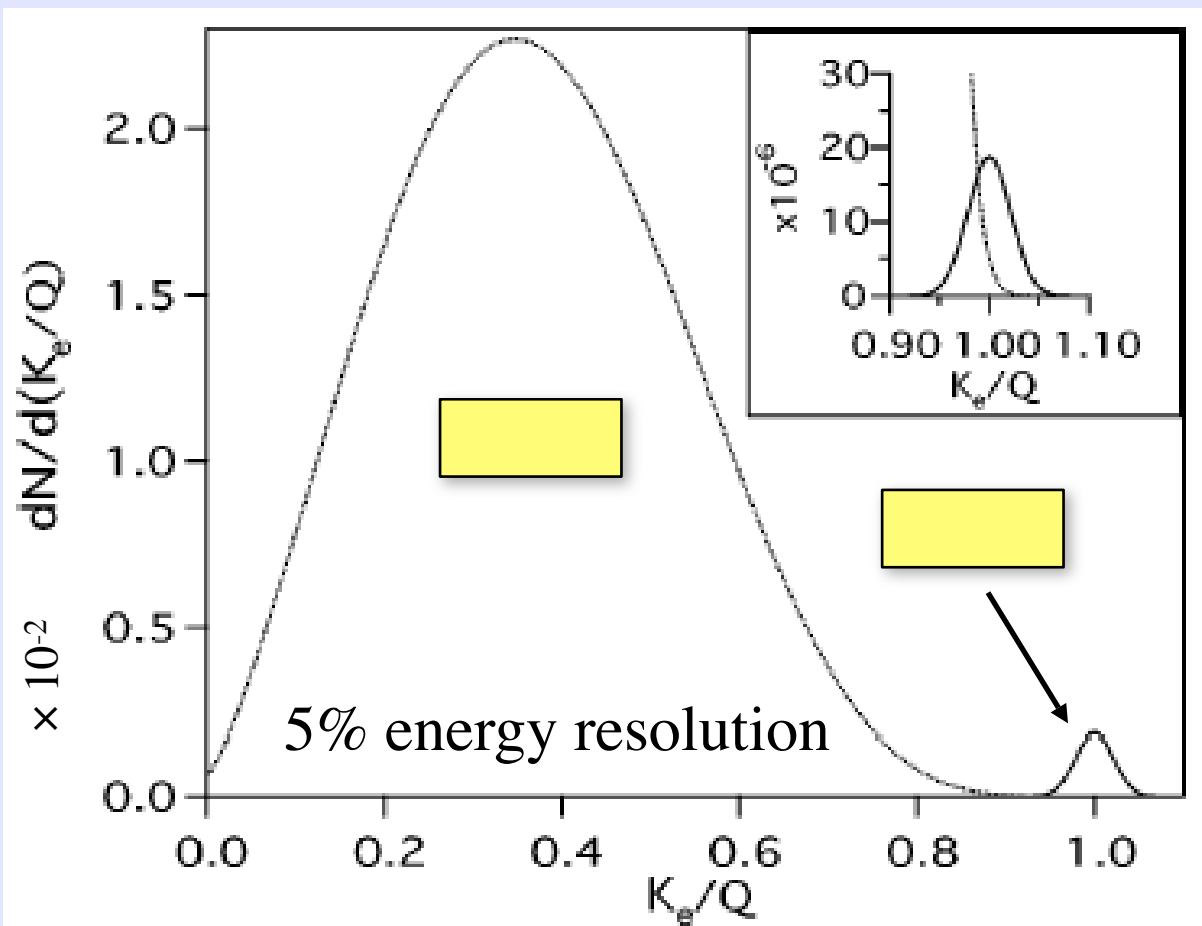


neutrinoless process ($0\nu\beta\beta$, $\Delta L=2$):



Observable for EXO

In the ideal case of perfectly efficient detection of the both the electrons and the Ba daughter, the only background to $0\nu\beta\beta$ is the spillage from the $2\nu\beta\beta$ channel due to finite energy resolution



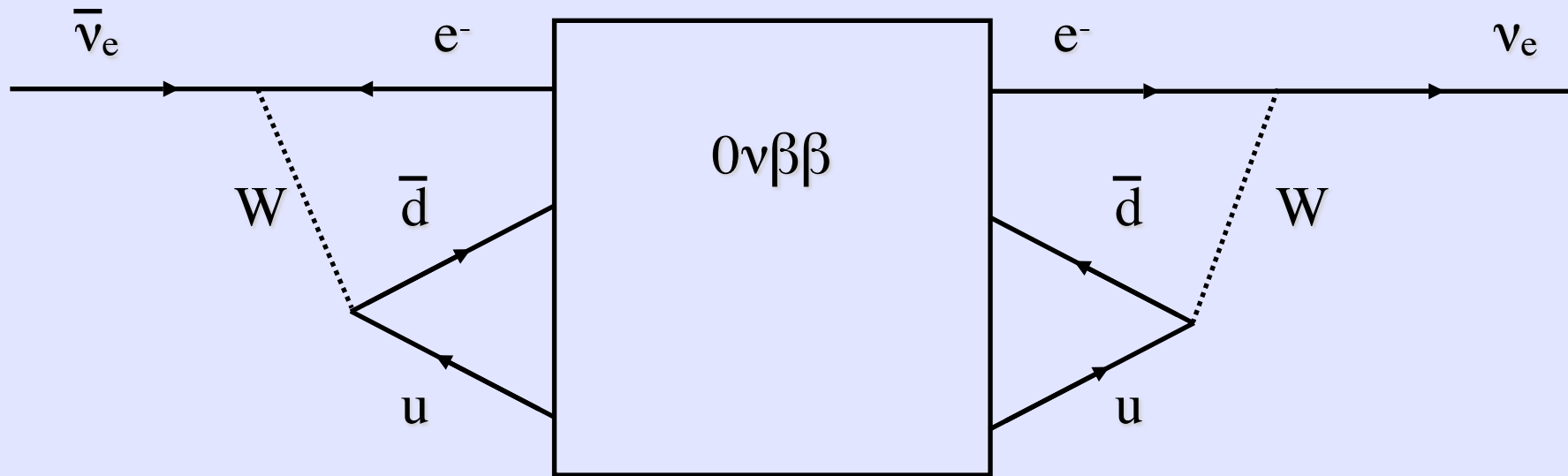
for gaussian profiles, leakage of $2\nu\beta\beta$ events into $0\nu\beta\beta$ peak goes like the 5.8th power of the energy resolution.

[Elliot and Vogel, Ann. Rev. Nucl. Part. Sci. 52 (2002) 115]

Majorana neutrinos

Neutrinoless double beta decay can only occur if neutrinos are their own antiparticles ($\nu = \bar{\nu}$):

[Schechter and Valle, Phys. Rev. D 25 (1982) 2951]



Neutrinoless double beta decay \Rightarrow neutrinos are Majorana particles regardless of the underlying mechanism

Why double beta decay?

If observed, neutrinoless double beta decay is the sign of new physics (although not completely unexpected new physics)

- Are neutrinos Majorana particles?
- What is the absolute mass scale of neutrinos?
- Other mechanisms? (complementary to other experiments)

Massive neutrinos

We know neutrinos oscillate between weak eigenstates, which happens because their weak and energy (mass) eigenstates do not coincide (i.e. the 3 mass eigenstates cannot be all zero):

$$|\nu_\alpha\rangle = \sum_i U_{\alpha i}^* |\nu_i\rangle$$

From oscillation experiments (solar, atmospheric, reactor, accelerator) we know only the **mass differences** between neutrino mass states:

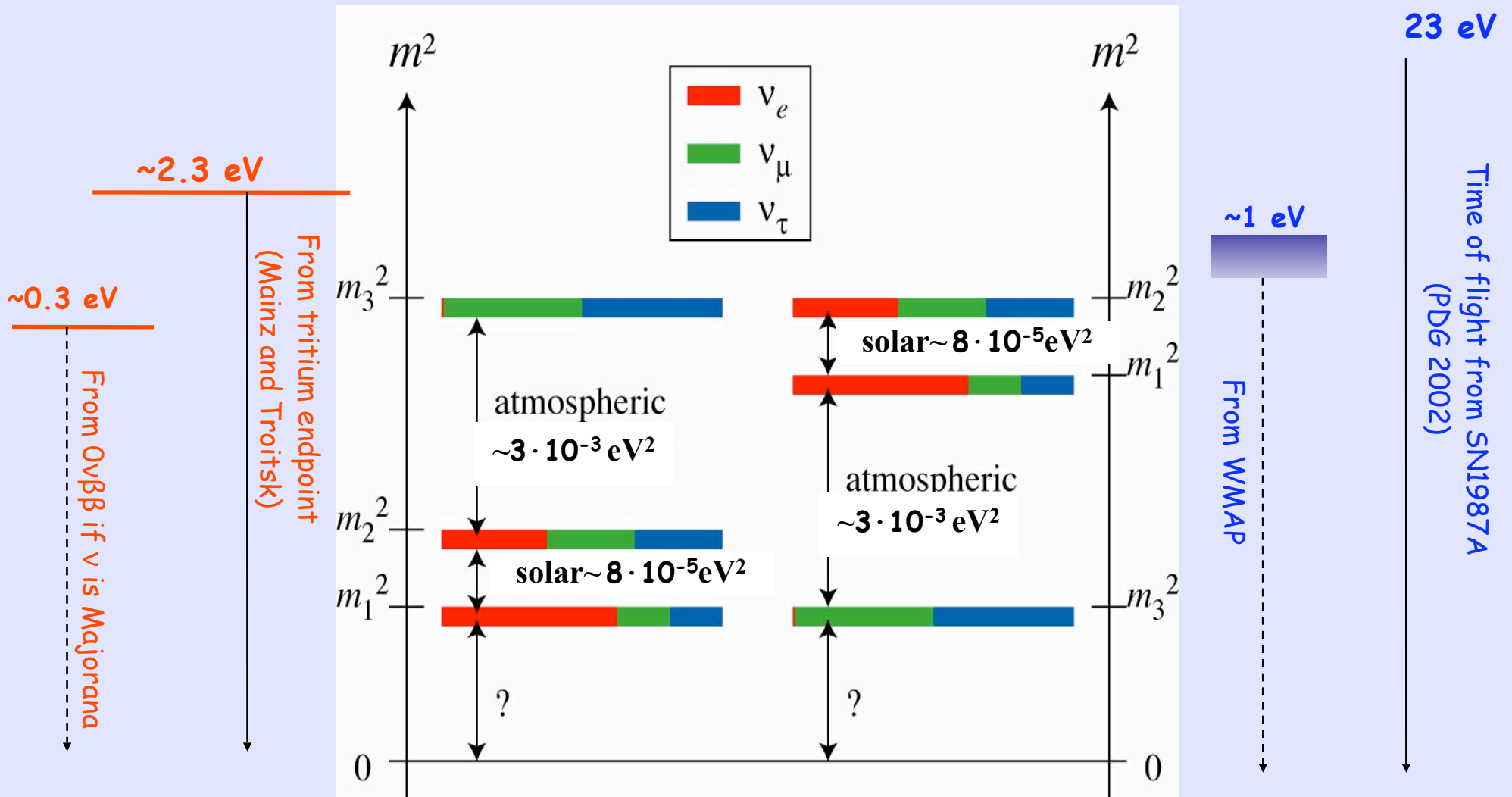
$$\begin{aligned} \Delta m_{12}^2 &\sim 8 \times 10^{-5} \text{ eV}^2 && \text{(solar neutrino experiments, KamLAND)} \\ \tan^2 \theta_{12} &\sim 0.4 \text{ (LMA-MSW)} \end{aligned}$$

$$\begin{aligned} \Delta m_{23}^2 &\sim 3 \times 10^{-3} \text{ eV}^2 && \text{(SuperK + Kamiokande, K2K, MINOS)} \\ \sin^2 \theta_{23} &> 0.9 \end{aligned}$$

$$\begin{aligned} \Delta m^2 &\sim 1 \text{ eV}^2 \\ \sin^2 \theta &\sim 3 \times 10^{-3} \\ &\text{(LSND, miniBooNE?)} \end{aligned}$$

Neutrino masses

We don't know if neutrino masses are degenerate (all very different from zero) and whether their hierarchy is direct or inverted



Neutrinos save lives!

Named for a subatomic particle with almost zero mass ...

Gear - Neutrino Carabiner - Netscape
View Go Communicator Help
Forward Reload Home Search Netscape Print
Location: http://www.mgear.com/mgear/itempg_3.icl

MOUNTAIN GEAR

YOUR ADVENTURE STARTS HERE

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Climbing ^ Rock Climbing ^ Carabiners



Neutrino Carabiner

by Black Diamond Equipment

Original Price: 8.50

Volume Discount: 6 for 7.83 each

Named for a subatomic particle with almost zero mass, this is the lightest, full-service carabiner made. That means it's the best choice for anyone who demands super lightweight carabiners without a compromise in strength. The mere 36 grams provide a large rope-bearing surface, a nose hood to protect against "gate rub", and a basket very similar to a Quicksilver 2.

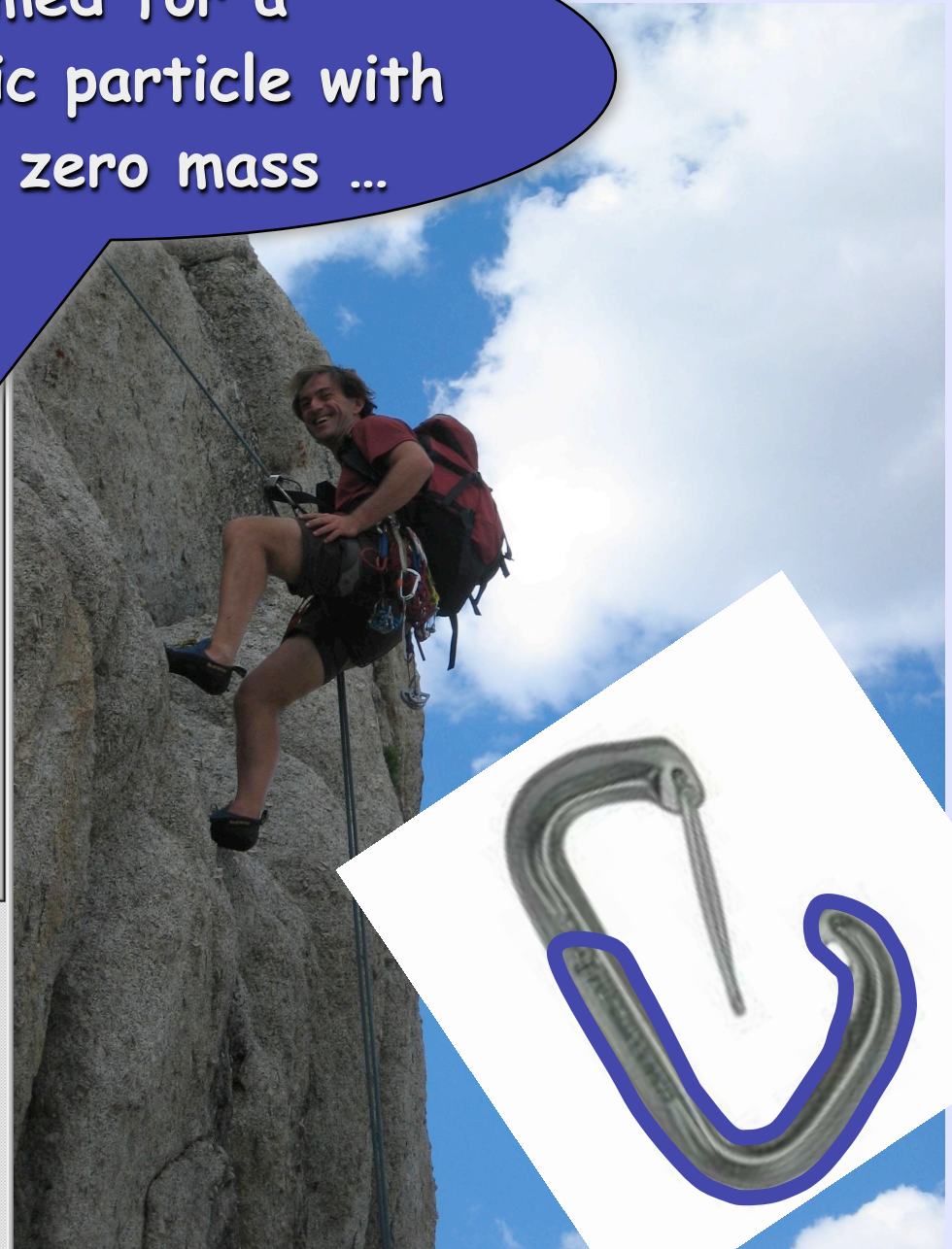
Prev

Next

QTY: 1

Add to Cart

Style	Weight	Strength	Strength (kN)		Gate Width
	grams	closed	open		(mm)
Neutrino	36	24	8		22



Double beta decay rate

$$\langle m_{\beta\beta} \rangle = \left(T_{1/2}^{0\nu\beta\beta} G^{0\nu\beta\beta}(Q, Z) |M_{nucl}|^2 \right)^{-1/2}$$

M_{nucl}

can be calculated within particular nuclear models ($\sim O(I)$)

$G^{0\nu\beta\beta}(Q, Z)$

a known phase space factor

$T_{1/2}^{0\nu\beta\beta}$

is the measured quantity [Hz]

$$\langle m_{\beta\beta} \rangle = \left| \sum_{i=1}^3 U_{ei}^2 m_i e^{i\alpha_i} \right|$$

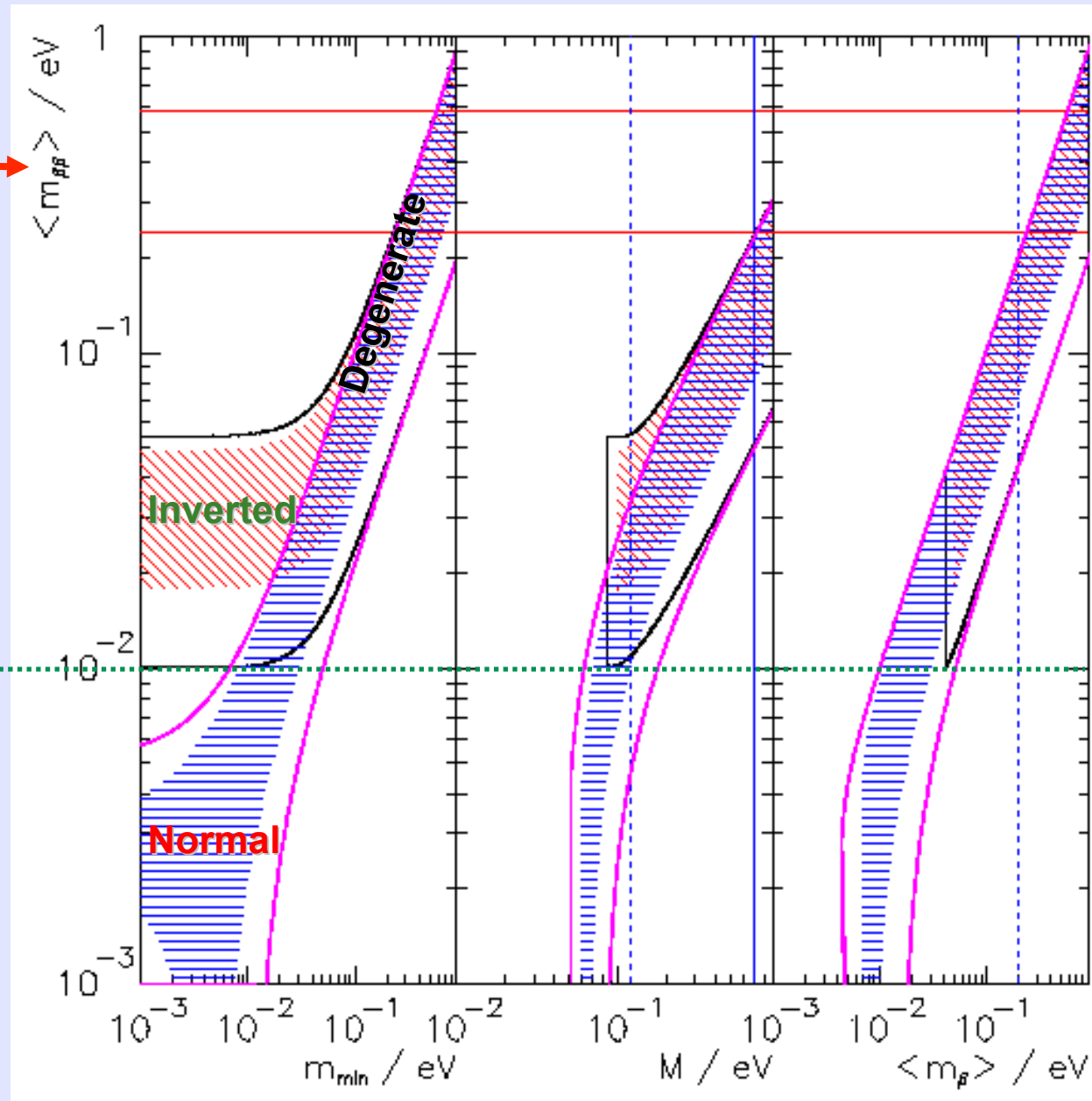
effective Majorana ν mass
($\epsilon_i = \pm 1$ if CP is conserved)

For reference, in direct kinematic searches of neutrino mass in β -decay:

$$\langle m_{\beta} \rangle^2 = \sum_i |U_{ei}|^2 m_i^2$$

Double beta decay effective neutrino mass

where we stand \rightarrow



^{76}Ge : 5.8×10^{28} years
 ^{100}Mo : 5.8×10^{28} years
 ^{136}Xe : 4.8×10^{28} years

[Rodin et al., Phys. Rev. C, 68 (2003)]

Double beta decay candidate isotopes

Candidate Q
(MeV) Abundance
(%)

$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	4.271	0.187
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	2.040	7.8
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	2.995	9.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	3.350	2.8
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	3.034	9.6
$^{110}\text{Pd} \rightarrow ^{110}\text{Cd}$	2.013	11.8
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	2.802	7.5
$^{124}\text{Sn} \rightarrow ^{124}\text{Te}$	2.228	5.64
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	2.533	34.5
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	2.459	8.9
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	3.367	5.6

$Q > 2 \text{ MeV}$:

- above most natural radioactive backgrounds
- higher rates
- need isotopic enrichment for all cases (except ^{130}Te)

Current limits on $0\nu\beta\beta$

Candidate nucleus	Detector type	(kg yr)	Present $T_{1/2}^{0\nu\beta\beta}$ (yr)	$\langle m \rangle$ (eV)
^{48}Ca			$>9.5 \cdot 10^{21}$ (76%CL)	(Phys. Lett. B 586 (2004) 198)
^{76}Ge	Ge diode	~30	$>1.9 \cdot 10^{25}$ (90%CL)	$<0.39^{+0.17}_{-0.28}$
^{82}Se			$>9.5 \cdot 10^{21}$ (90%CL)	
^{100}Mo			$>5.5 \cdot 10^{22}$ (90%CL)	
^{116}Cd			$>7.0 \cdot 10^{22}$ (90%CL)	
^{128}Te	TeO ₂ cryo	~3	$>1.1 \cdot 10^{23}$ (90%CL)	
^{130}Te	TeO ₂ cryo	~3	$>2.1 \cdot 10^{23}$ (90%CL)	$<1.1 - 2.6$
^{136}Xe	Xe scint	~10	$>1.2 \cdot 10^{24}$ (90%CL)	<2.9
^{150}Nd			$>1.2 \cdot 10^{21}$ (90%CL)	
^{160}Gd			$>1.3 \cdot 10^{21}$ (90%CL)	

Adapted from the Particle Data Group 2003

EXO strategy

Goal: ton-scale enriched Xe (80% ^{136}Xe) time projection chamber (TPC) with scintillation light collection and $^{136}\text{Ba}^+$ identification (possibility unique to ^{136}Xe)

Phased approach:

- EXO-200 “prototype” detector (200 kg of enriched xenon, 80% ^{136}Xe , no Ba tagging)
- Ba identification R&D as a parallel effort
- merge into one proposal

EXO-200

EXO-200 is a LXe TPC with scintillation light readout that employs 200 kg of enriched xenon (80% ^{136}Xe)

→ EXO-200 has no $^{136}\text{Ba}^+$ identification ←

Goals:

- look for $0\nu\beta\beta$ decay of ^{136}Xe with competitive sensitivity and test backgrounds of large LXe detector at ~ 2000 m.w.e. depth
($T_{1/2}^{0\nu} > 6 \times 10^{25}$ y, current limit: $T_{1/2}^{0\nu} > 1.2 \times 10^{24}$ y)
- measure the standard $2\nu\beta\beta$ decay of ^{136}Xe ($Q = 2457.8 \pm 0.4$ keV) and measure its lifetime (best upper limit to date: $T_{1/2}^{2\nu} > 1 \times 10^{22}$ y)
[\[R. Bernabei et al., Phys. Lett. B 546 \(2002\) 23\]](#)
- test LXe technology and enrichment on a large scale
- test TPC components, light readout (518 LAAPDs), and radioactivity of materials, xenon handling and purification

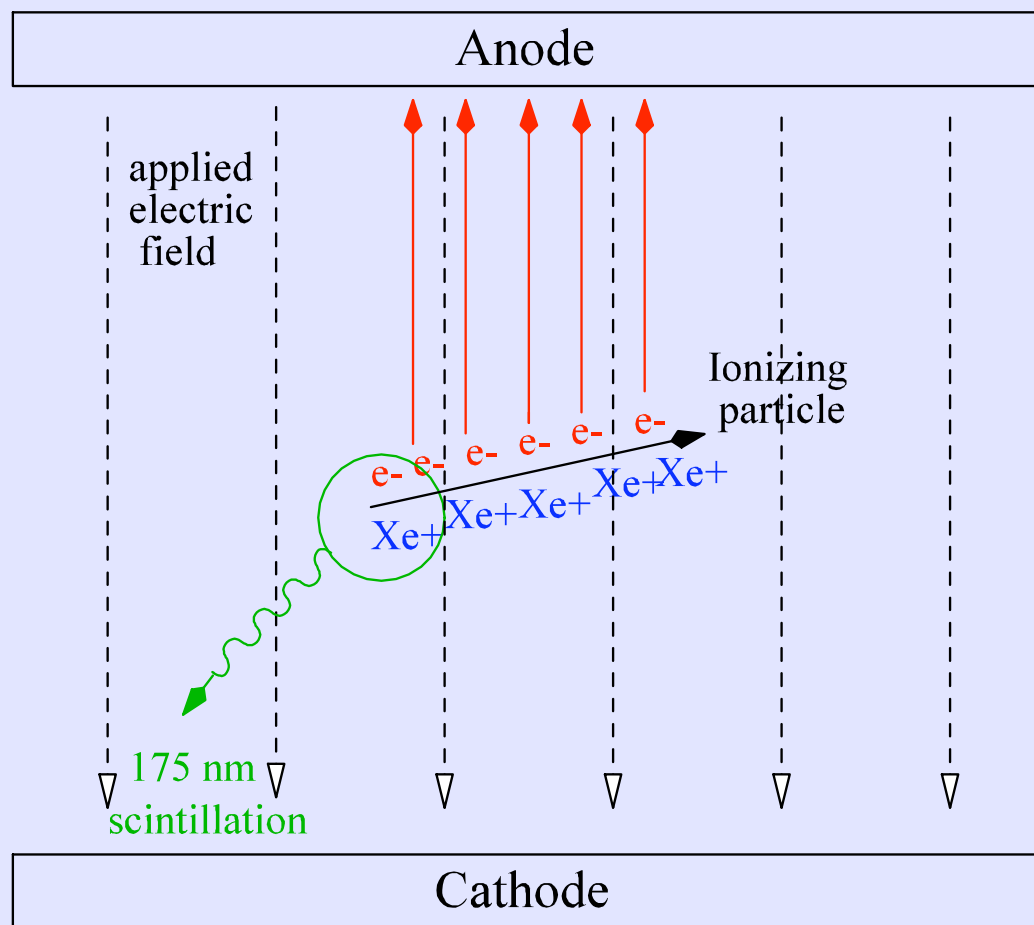
2νββ event rate

2νββ decay has never been observed in ^{136}Xe . Some of the lower limits on its half life are close to (and in one case below) the theoretical expectation.

	$T_{1/2}$ (yr)	evts/year in the 200kg prototype (no efficiency applied)
Experimental limit		
Leuscher et al	$>3.6 \cdot 10^{20}$	$<1.3 \text{ M}$
Gavriljuk et al	$>8.1 \cdot 10^{20}$	$<0.6 \text{ M}$
Bernabei et al	$>1.0 \cdot 10^{22}$	$<48 \text{ k}$
Theoretical prediction		
QRPA (Staudt et al.) [$T_{1/2}^{\text{max}}$]	$=2.1 \cdot 10^{22}$	$=23 \text{ k}$
QRPA (Vogel et al.)	$=8.4 \cdot 10^{20}$	$=0.58 \text{ M}$
NSM (Caurier et al.)	$=2.1 \cdot 10^{21}$	$=0.23 \text{ M}$

EXO-200 is very well positioned to solve this issue
(67 decays/day/100 kg)

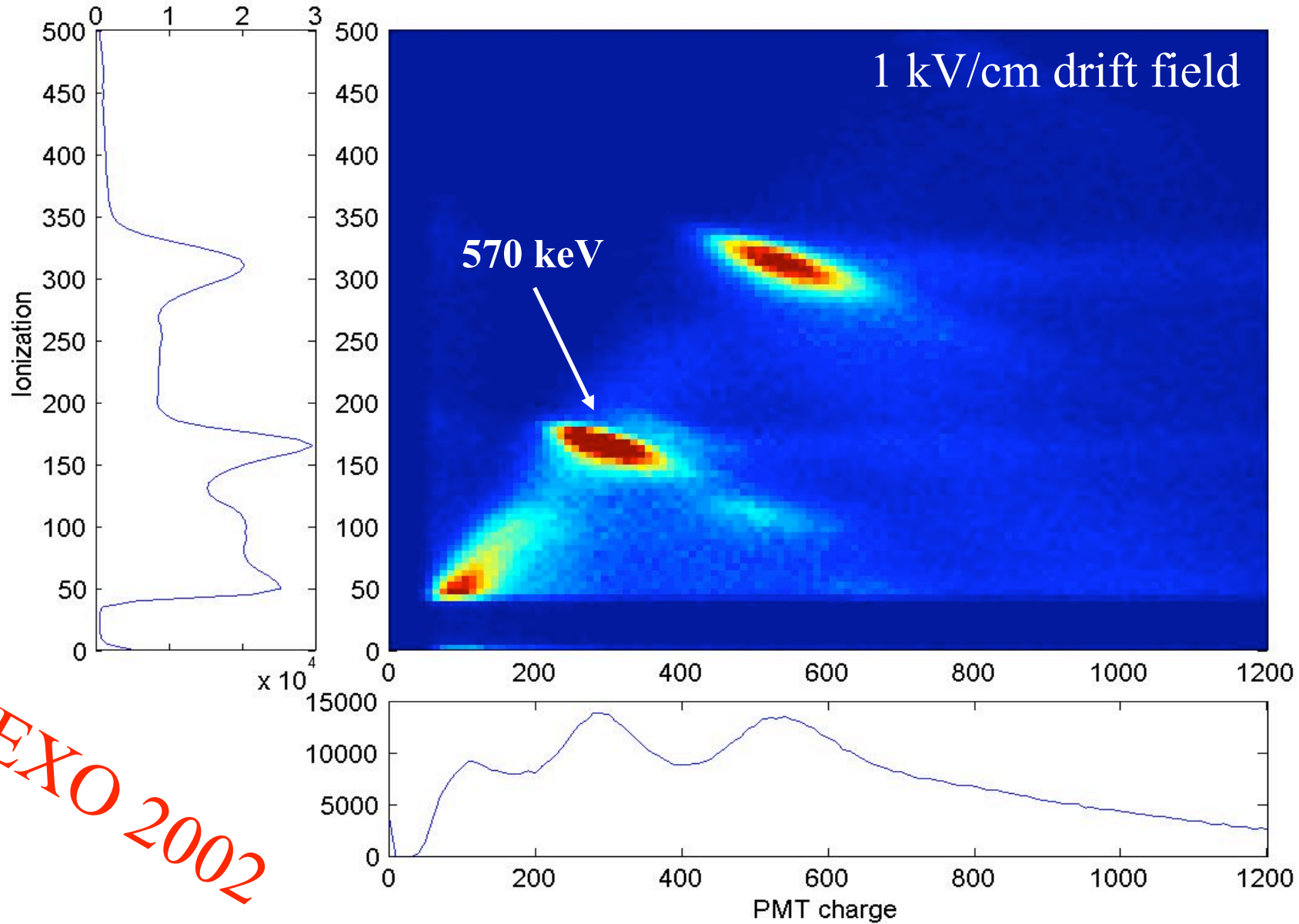
Dual readout: ionization and scintillation



The position of the event is reconstructed with two sets of wires at the anode (x-y) and the drift time (z). The position is important for external background identification and, in the future, for Ba ion identification

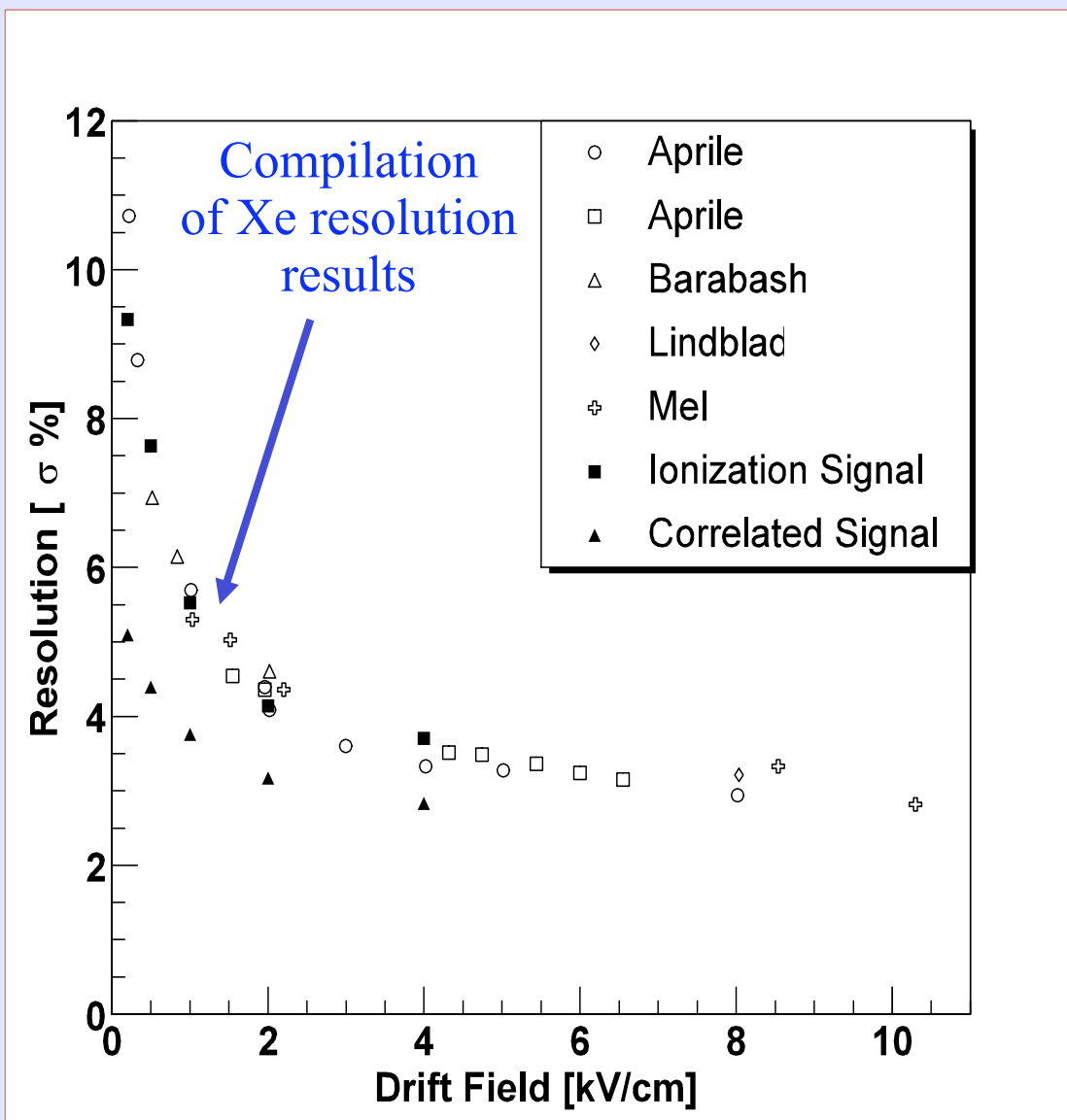
The event energy is measured by collecting the ionization charge on the anode and/or the amplitude of the scintillation light pulse (used as $t=0$ reference for the z drift).

Data show microscopic anticorrelation between ionization and scintillation



EXO 2002

Anti-correlated ionization and scintillation improves the energy resolution in LXe



[E. Conti et al., Phys. Rev. B: 68 054201]

Ionization alone:

$$\sigma(E)/E = 3.8\% \text{ @ } 570 \text{ keV}$$

or 1.8% @ $Q_{\beta\beta}$

Ionization & Scintillation:

$$\sigma(E)/E = 3.0\% \text{ @ } 570 \text{ keV}$$

or 1.4% @ $Q_{\beta\beta}$

(twice as good as most recent xenon $\beta\beta 0\nu$ experiment)

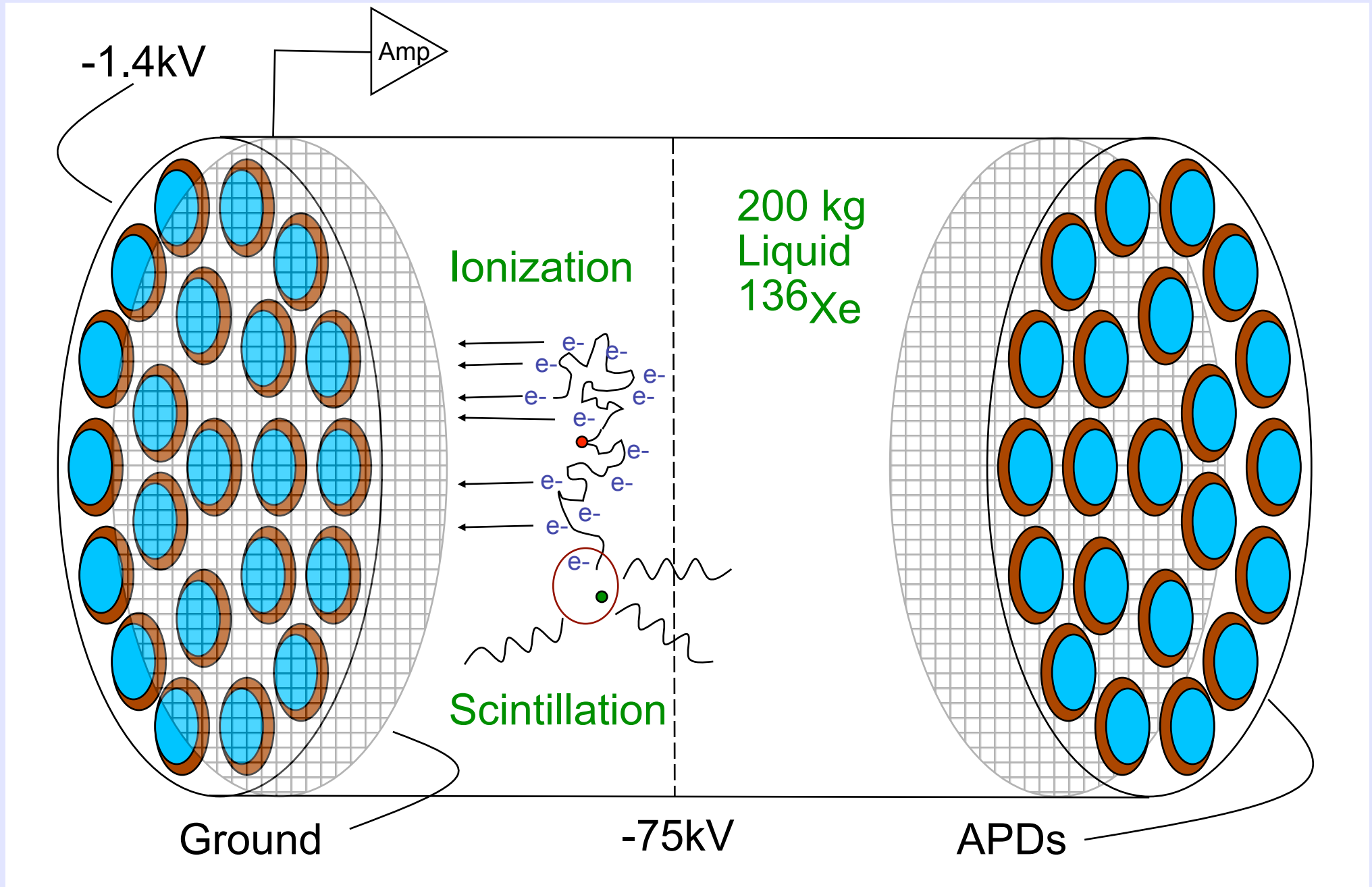
EXO-200 will collect 3-4 times as much scintillation... possibly giving further improvement (one of the main technical tasks to be addressed)

Resolution improvement is very important to separate the $0\nu\beta\beta$ and $2\nu\beta\beta$ modes

The EXO-200 TPC

- Two symmetric drift regions (16.5 cm long) along the cylinder axis, defined by a central cathode plane running at negative high voltage
 - max high voltage is 70 kV (3.5 kV/cm drift field); energy resolution improves with drift field, but possibly lower fields will allow for a better separation between events with 1 and 2 electrons (optimization is part of EXO-200's goals)
 - two sets of crossed anode wires (3 mm pitch, 100 μ m diameter) at each end of the cylinder, read out in groups of 3 (48 \times 48 channels), for a total of 96 channels per $\frac{1}{2}$ detector
- 259 Large Area Avalanche Photodiodes (LAAPDs) at each end of the cylinder, behind the anode wires (90% light transmission)
 - "bare" devices, DUV sensitive (QE \sim 1 @ 175 nm)
- y-position given by induction signal on shielding grid.
x-position and energy given by charge collection grid.
APD array observes prompt scintillation to measure drift time and improve energy resolution.

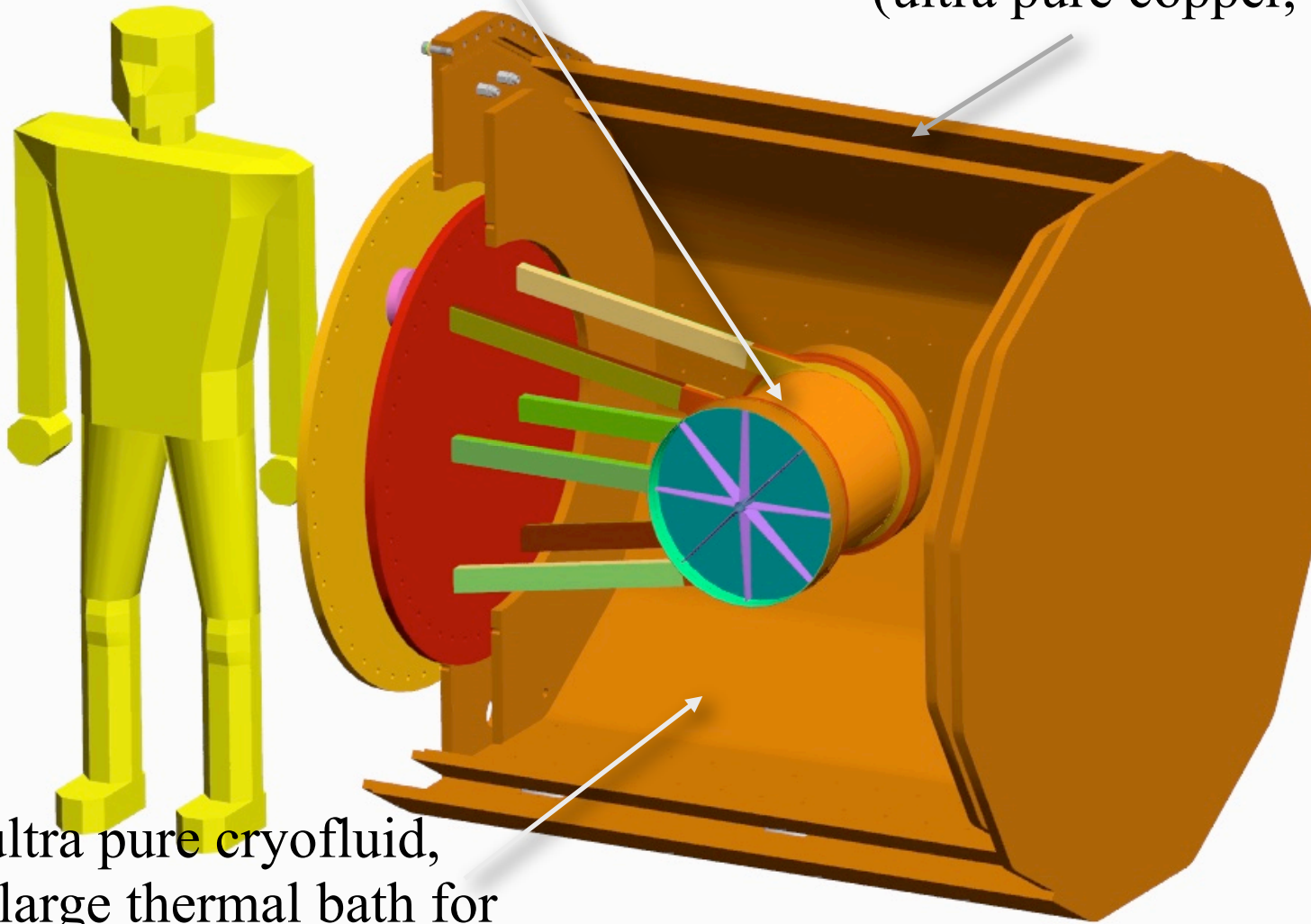
EXO-200 detector schematic



The EXO-200 detector

200 kg of LXe in thin vessel
(ultra pure copper, 1.5 mm thick)

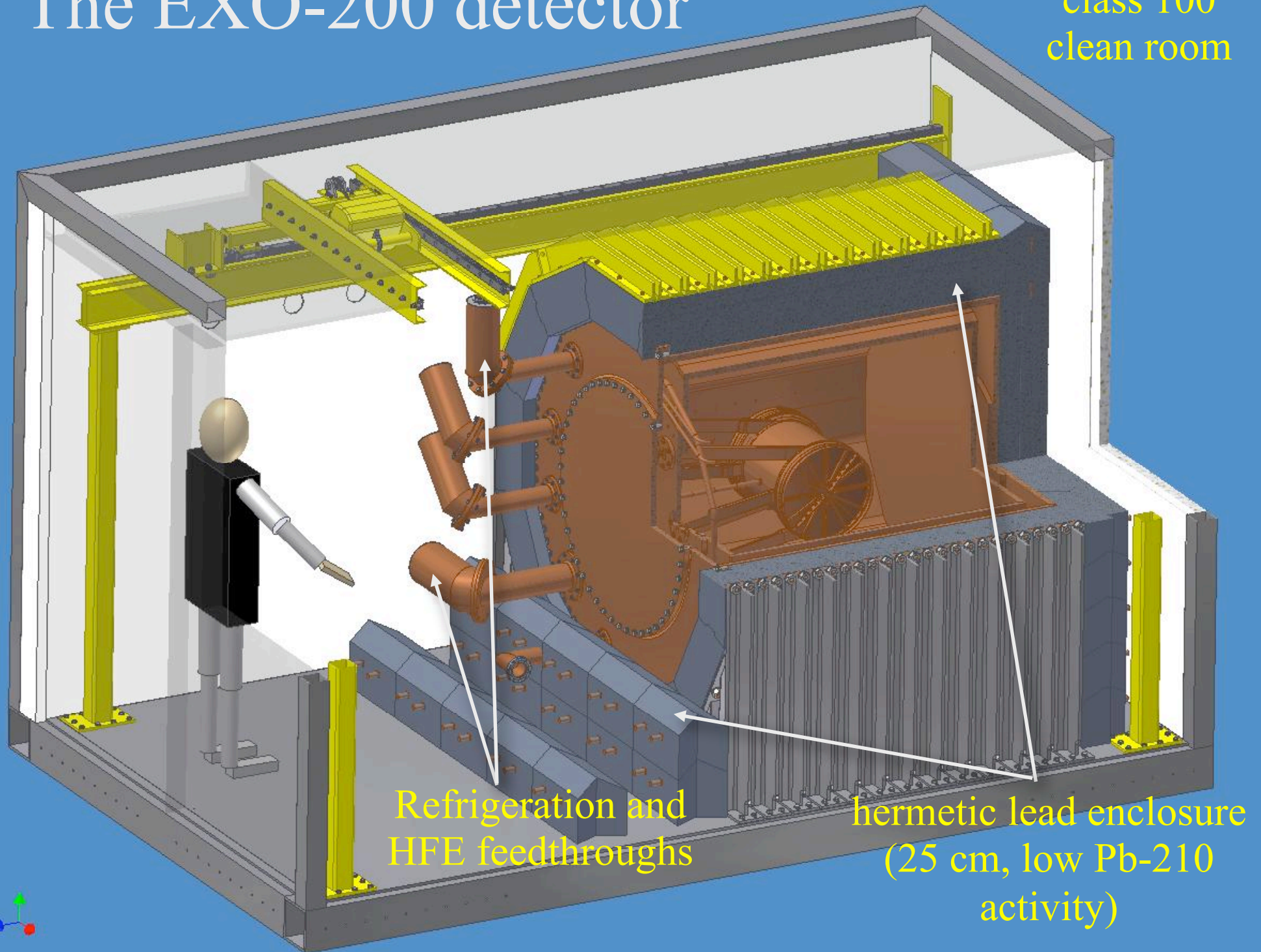
double walled vacuum insulated
cryostat
(ultra pure copper, 2.5 cm thick)



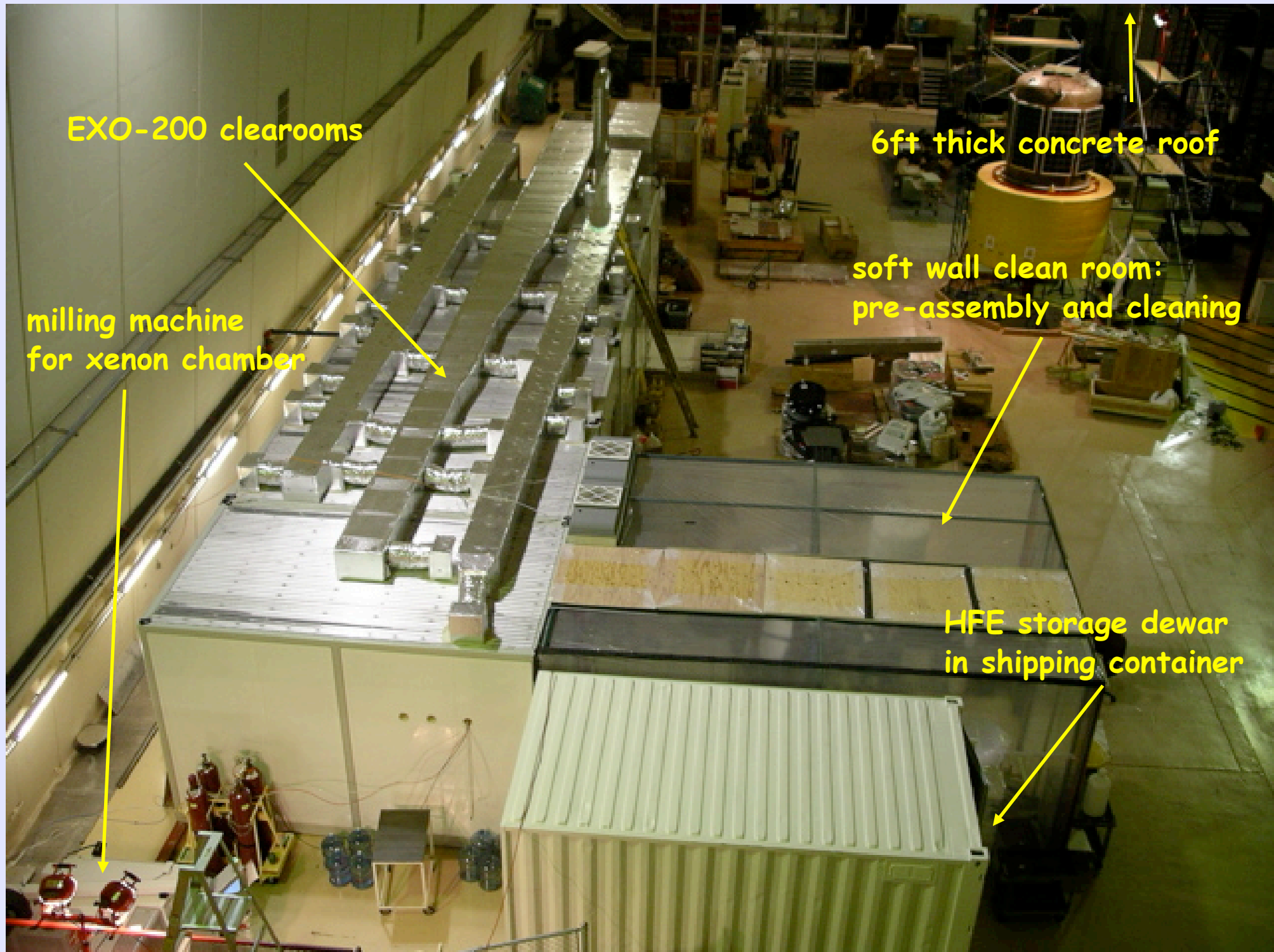
50 cm of ultra pure cryofluid,
providing large thermal bath for
uniform temperature
(3M HFE-7000, hydrofluoroether $C_3F_7OCH_3$)

The EXO-200 detector

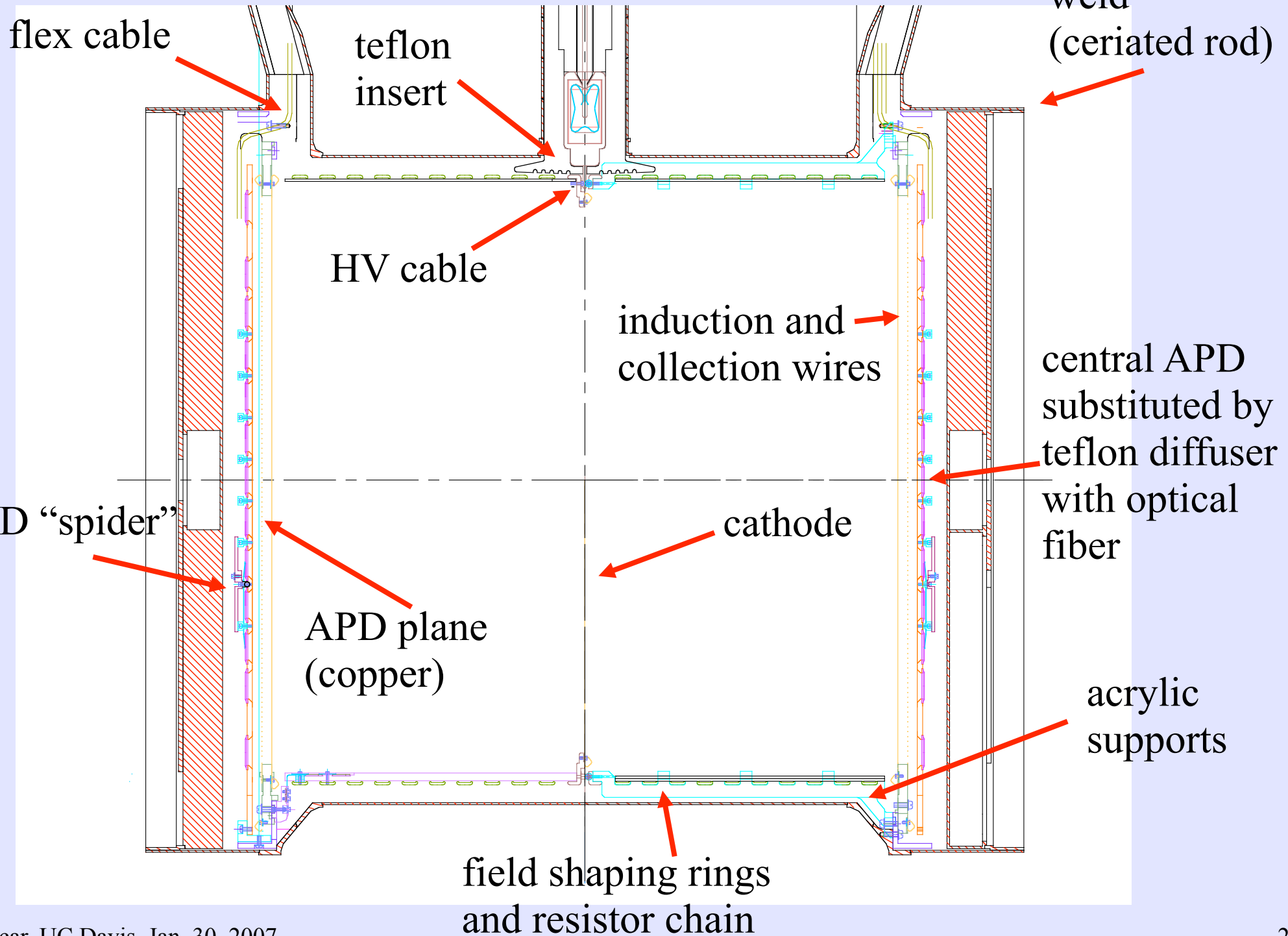
class 100
clean room



The EXO-200 modular clean rooms

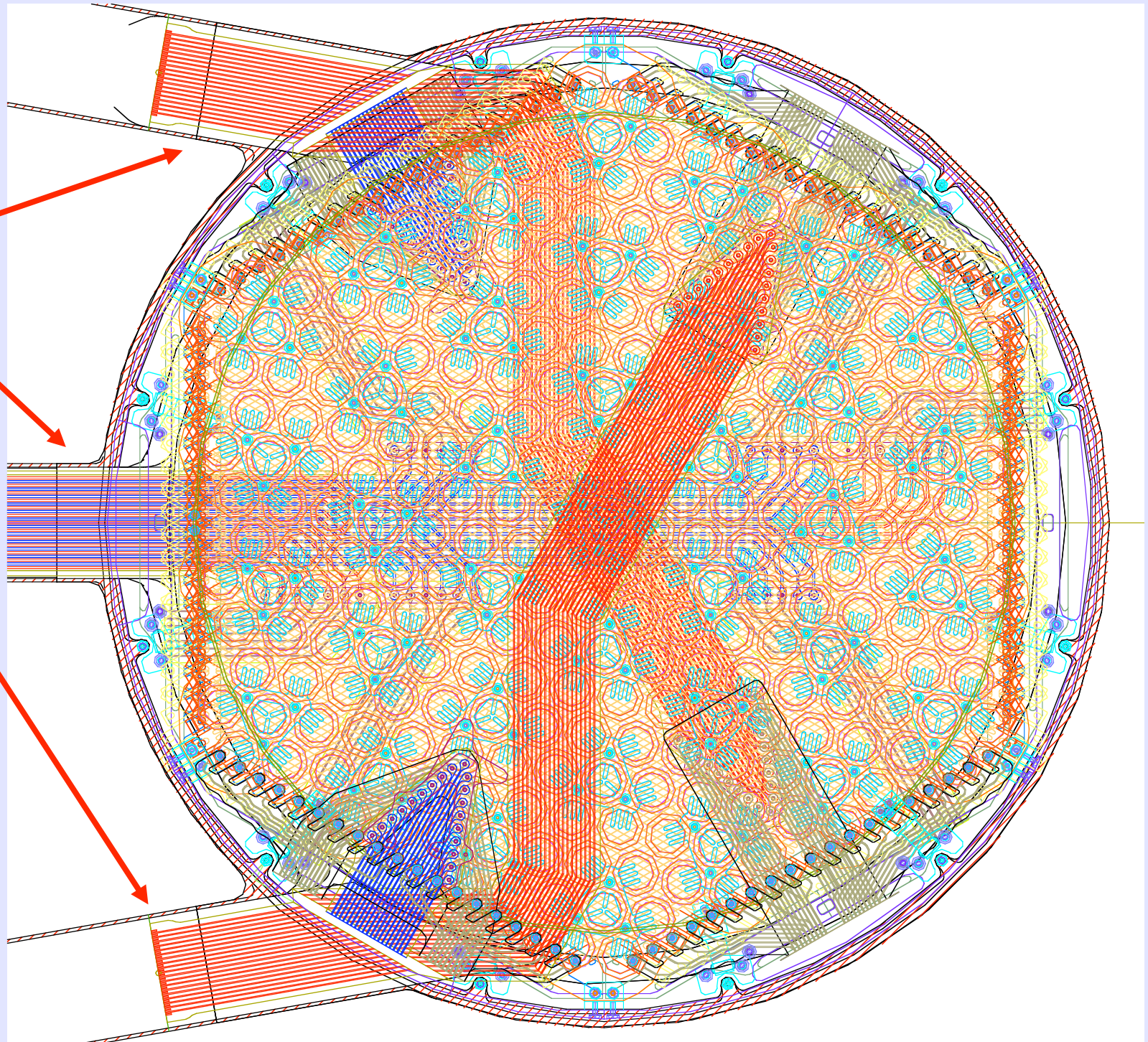


EXO-200 TPC



EXO-200 TPC

stripline cables
routed through
the legs



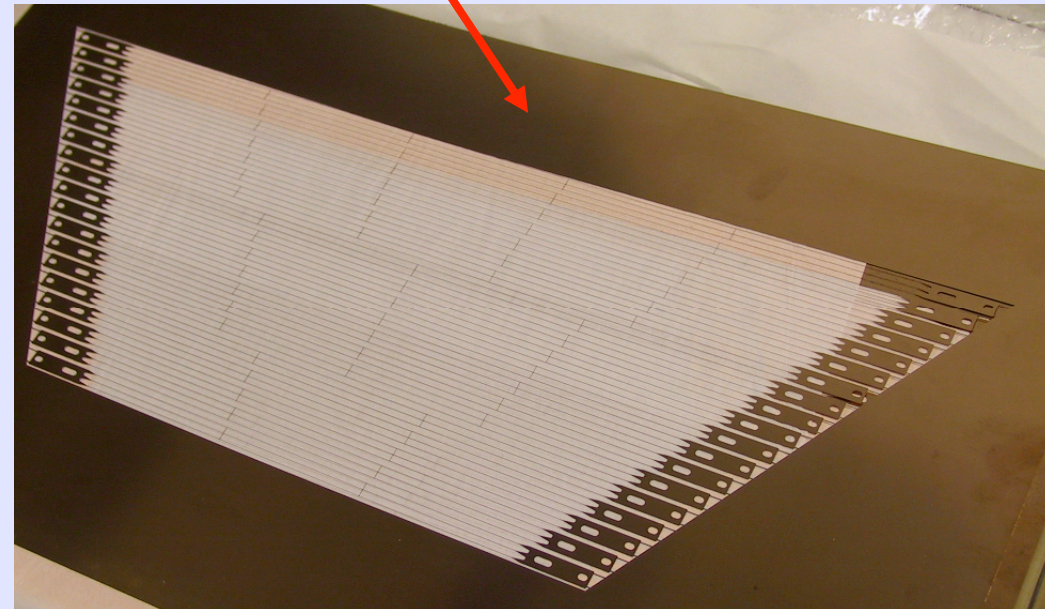
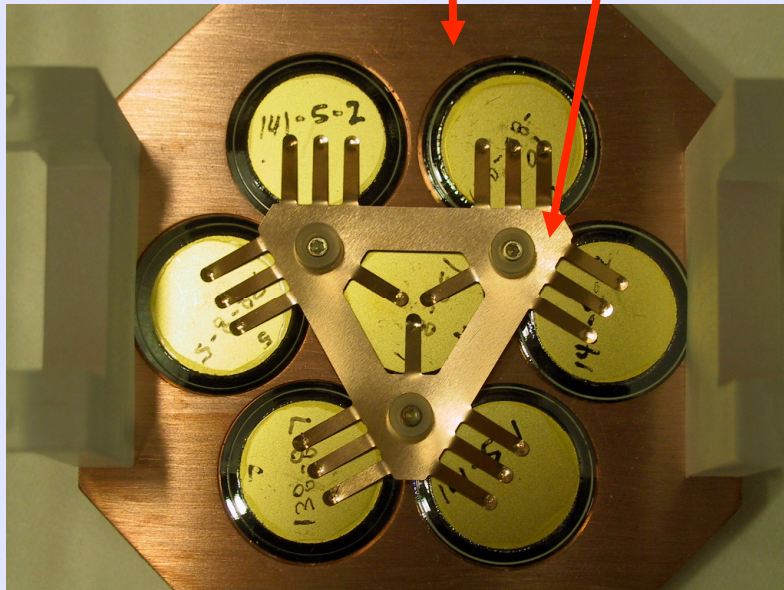
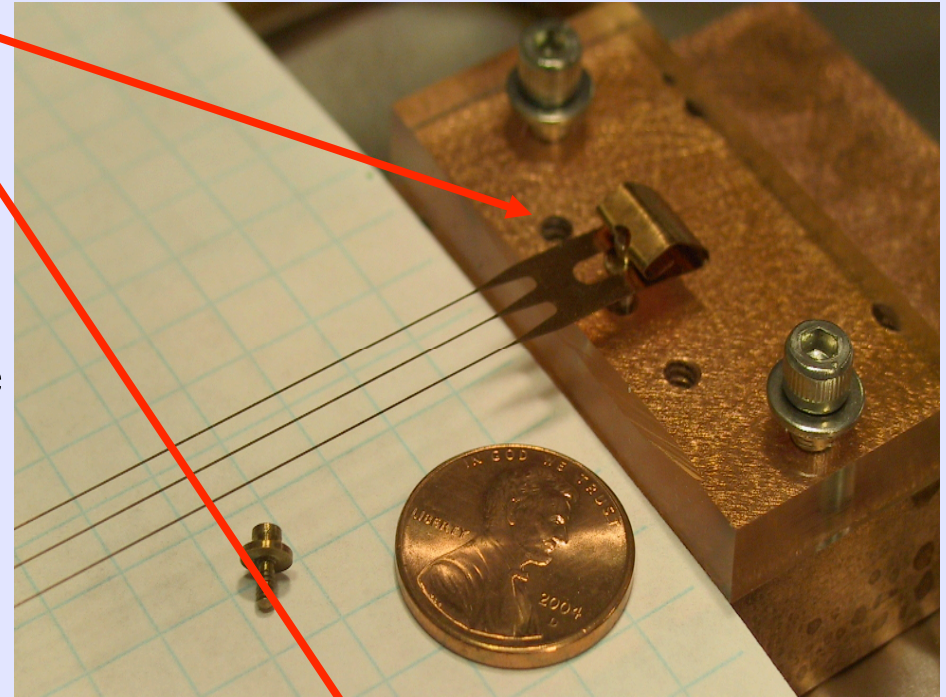
EXO-200 TPC

TPC fully designed, including:

- photoetched wires (3 wires / channel)
- photoetched cathode plane
- photoetched “spider” holders for APDs
- APD plane (groups of 7)
- teflon reflector (from DuPont TE-6472)
- field grading resistor chain
- stripline cable layout and connection scheme

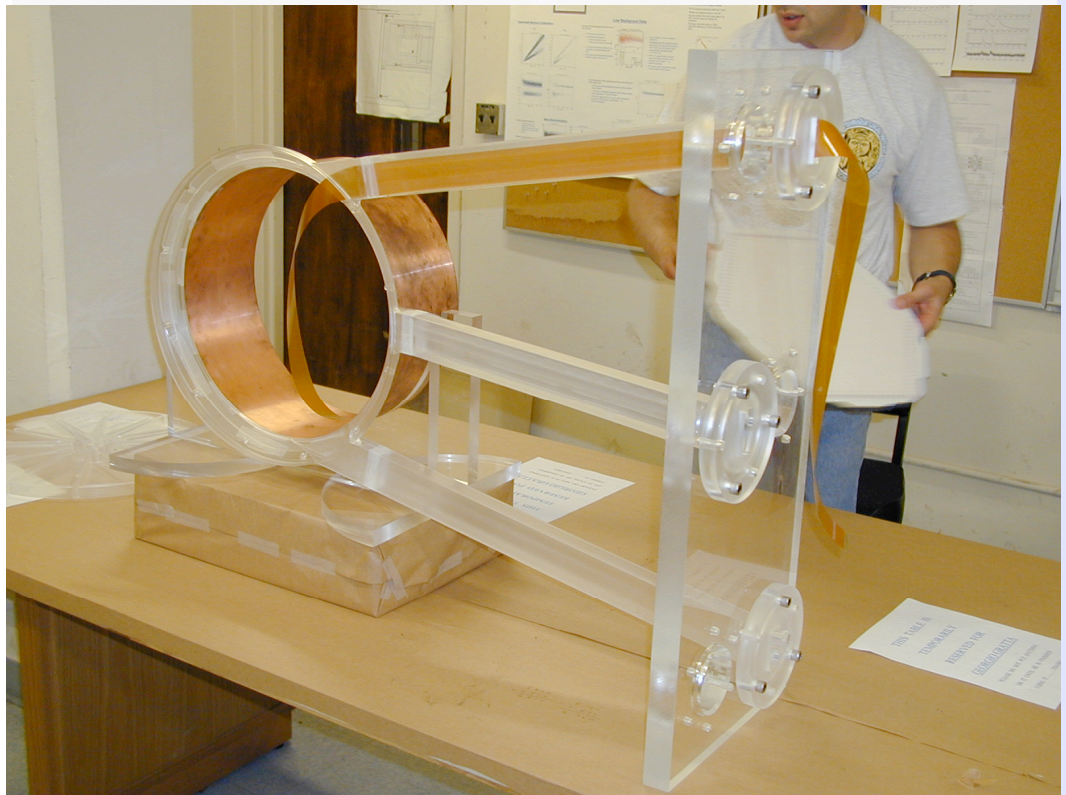
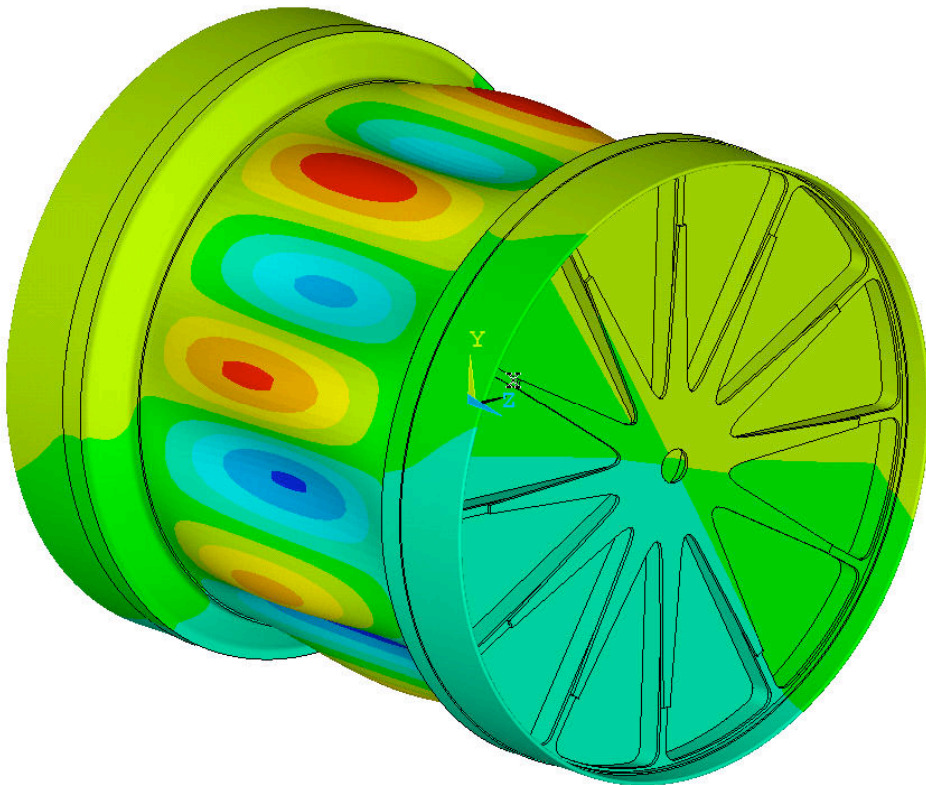
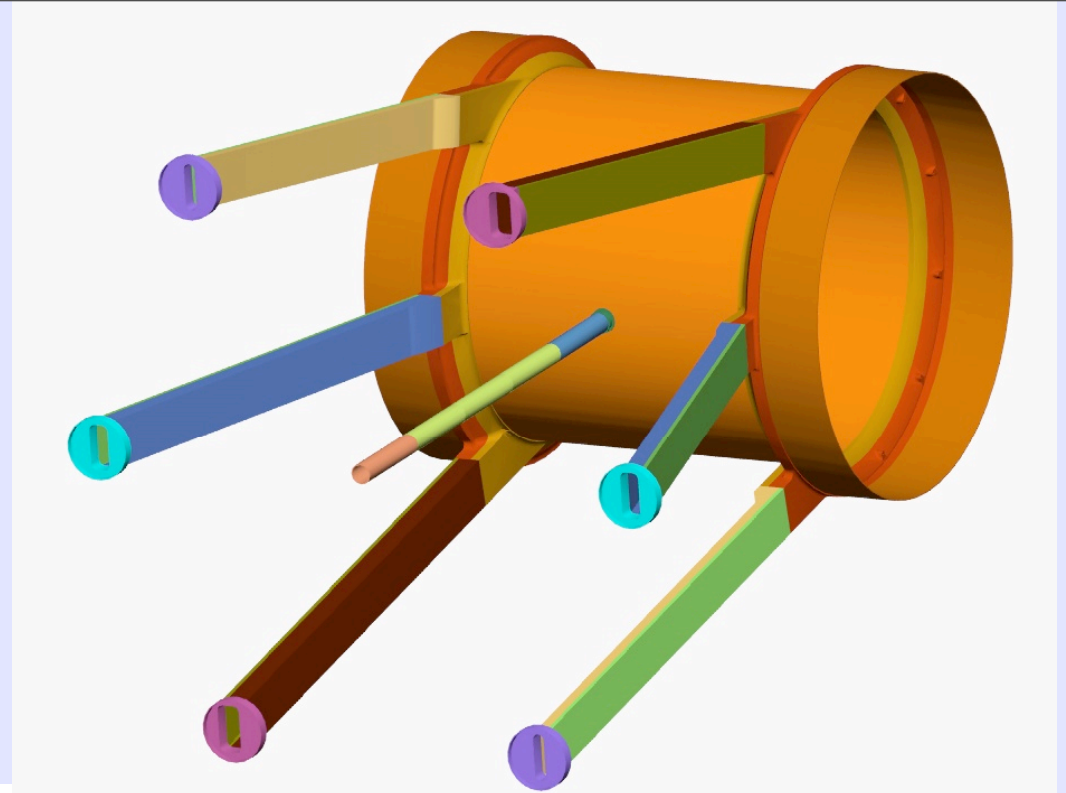
Approved material at hand for:

- photoetched parts (produced)
- resistors, silicon bronze screws, epoxy
- teflon reflectors, stripline cables

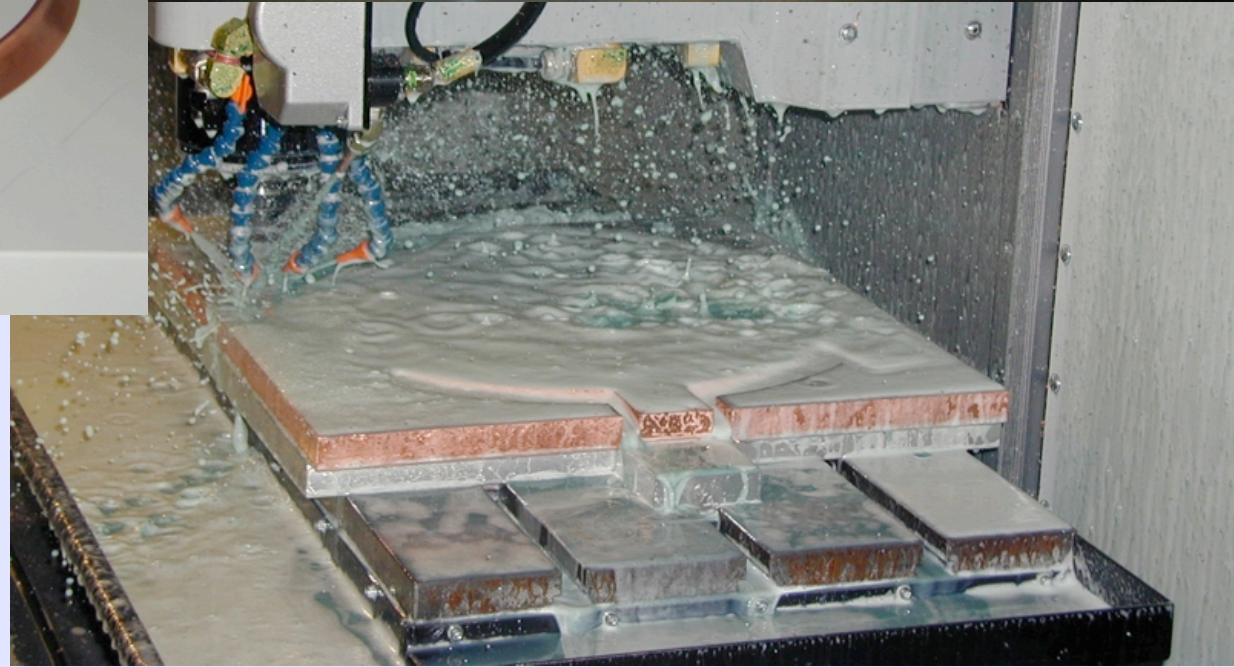


EXO-200 xenon vessel

- drawings and FEA structural analysis complete
- full scale model of half detector with legs produced
- all machining performed in ES3 at Stanford (2 m concrete)
- e-beam welding performed in East Bay - 1 day exposure



EXO-200 xenon vessel machining

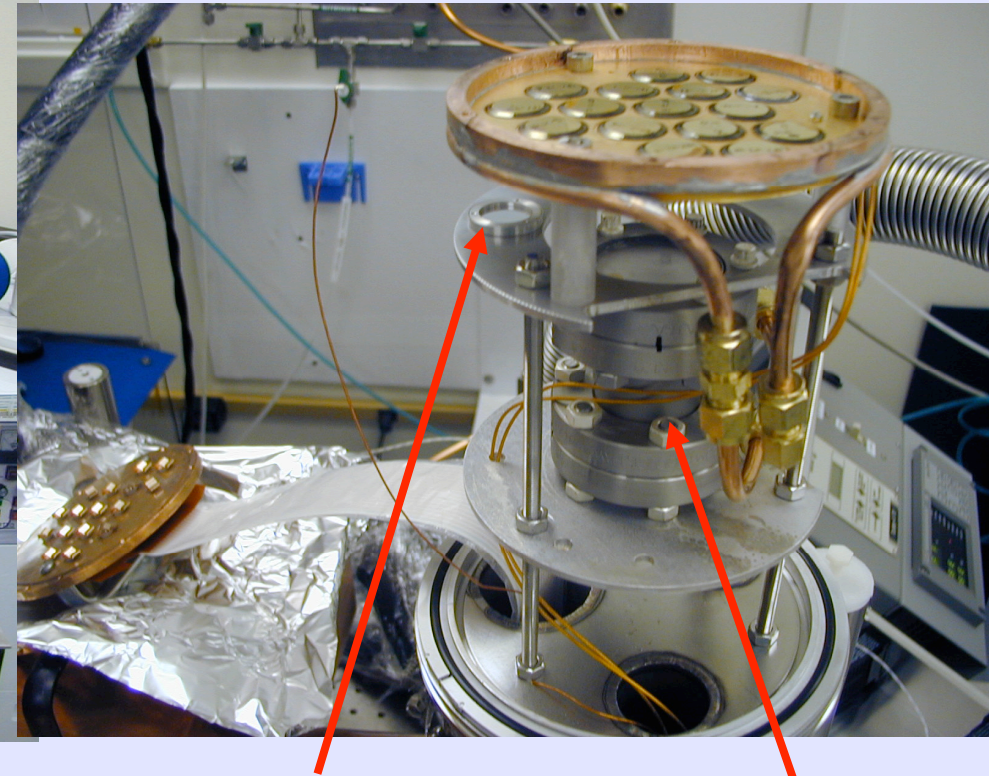
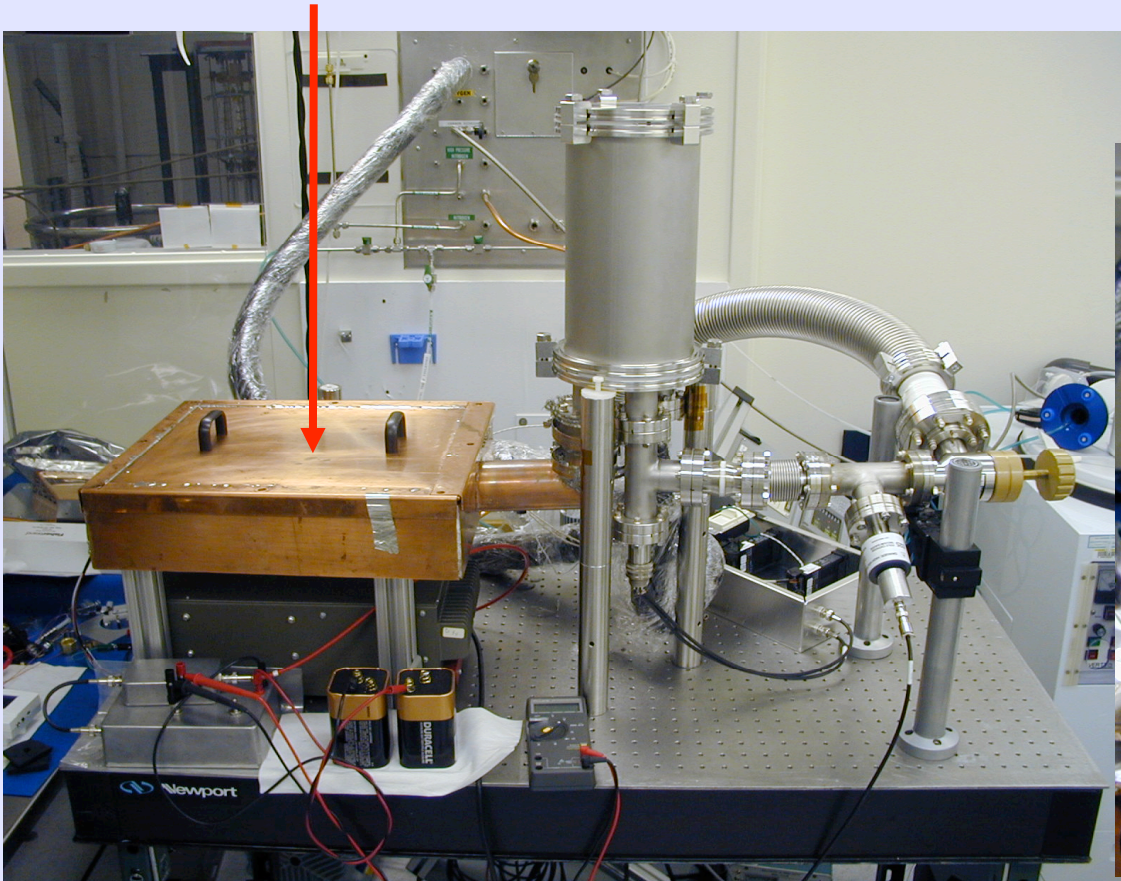


- machining of parts has begun last week

EXO-200 readout board

LAAPD testing rig

16 APDs tested together



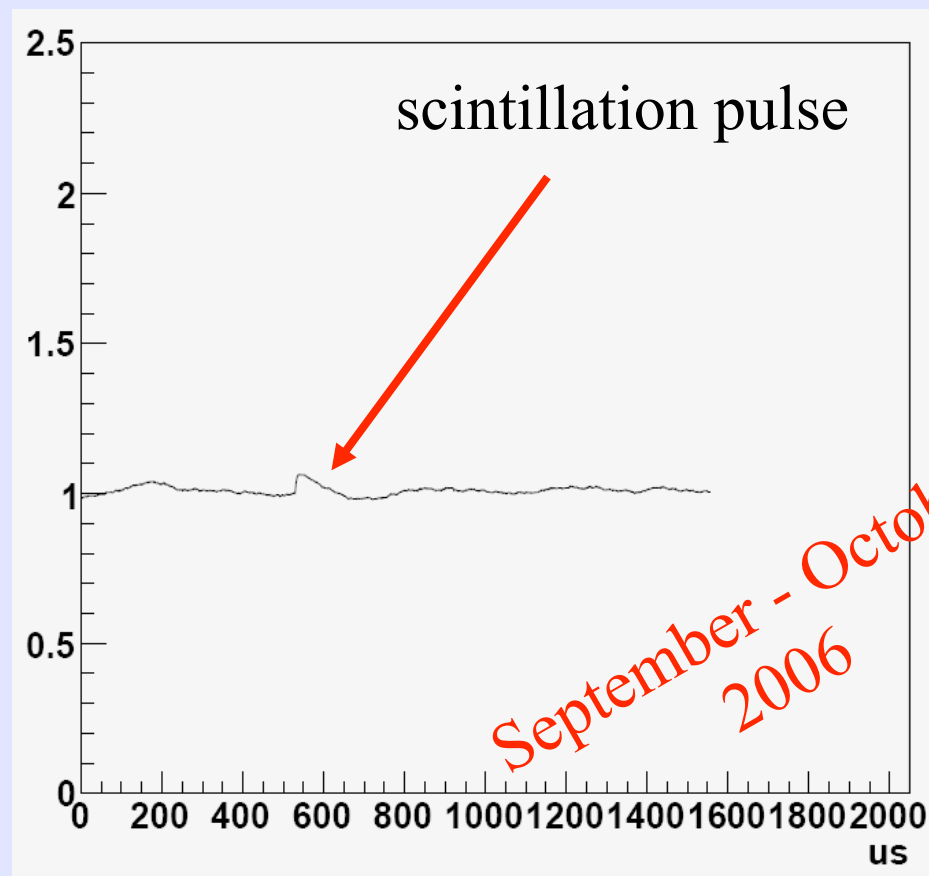
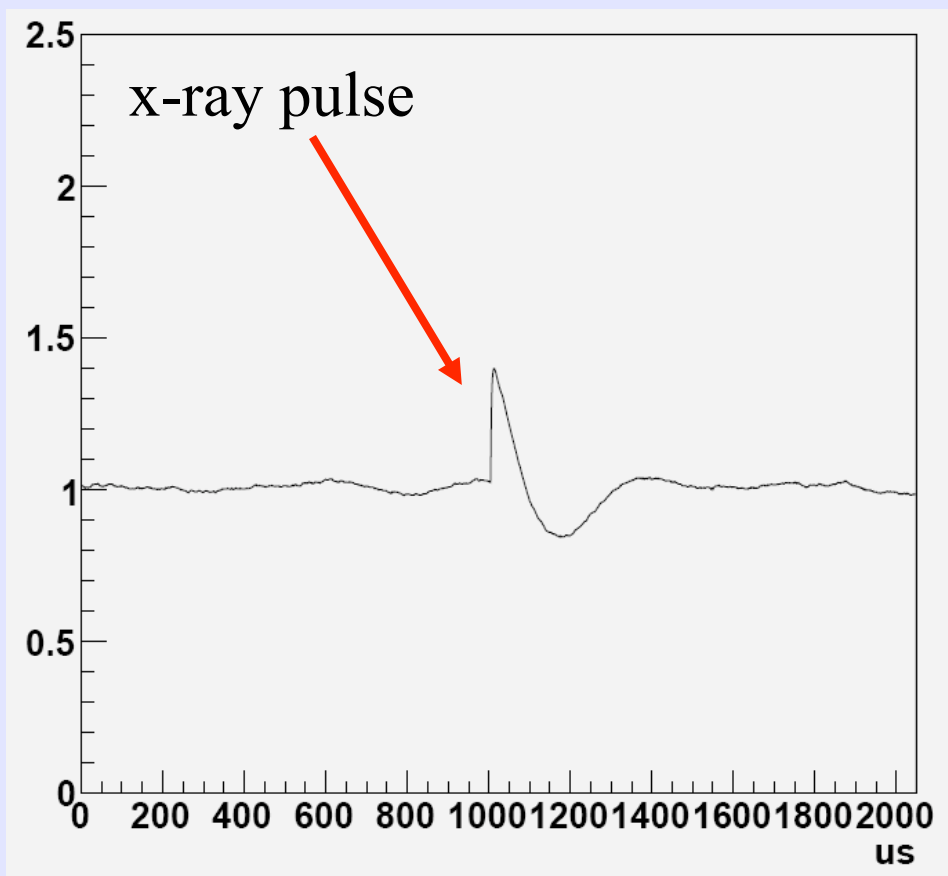
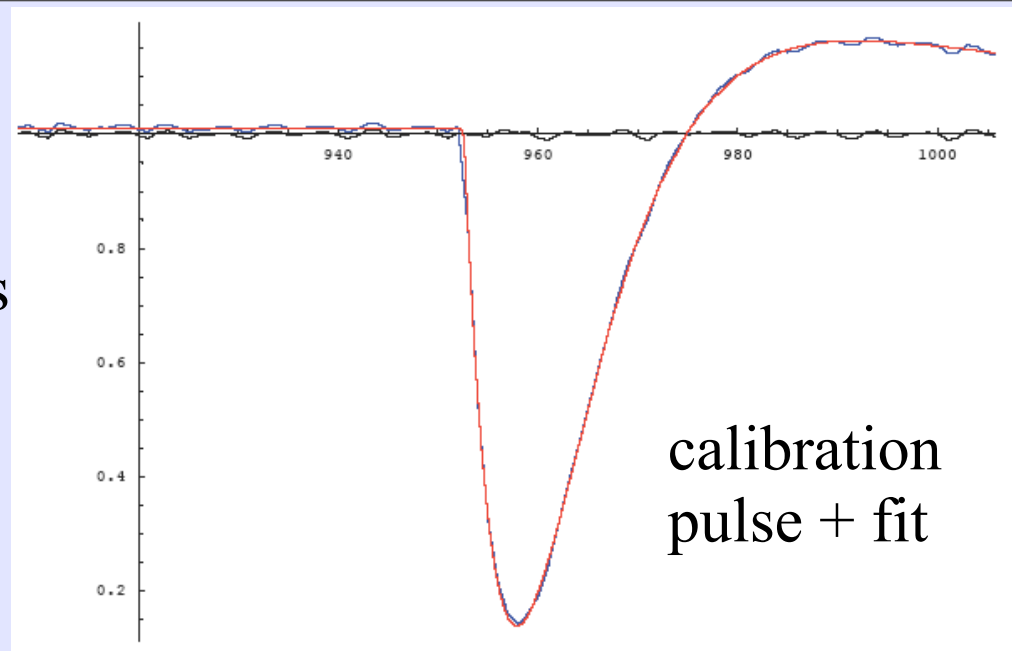
Fe-55 x-ray source

xenon
scintillation
source

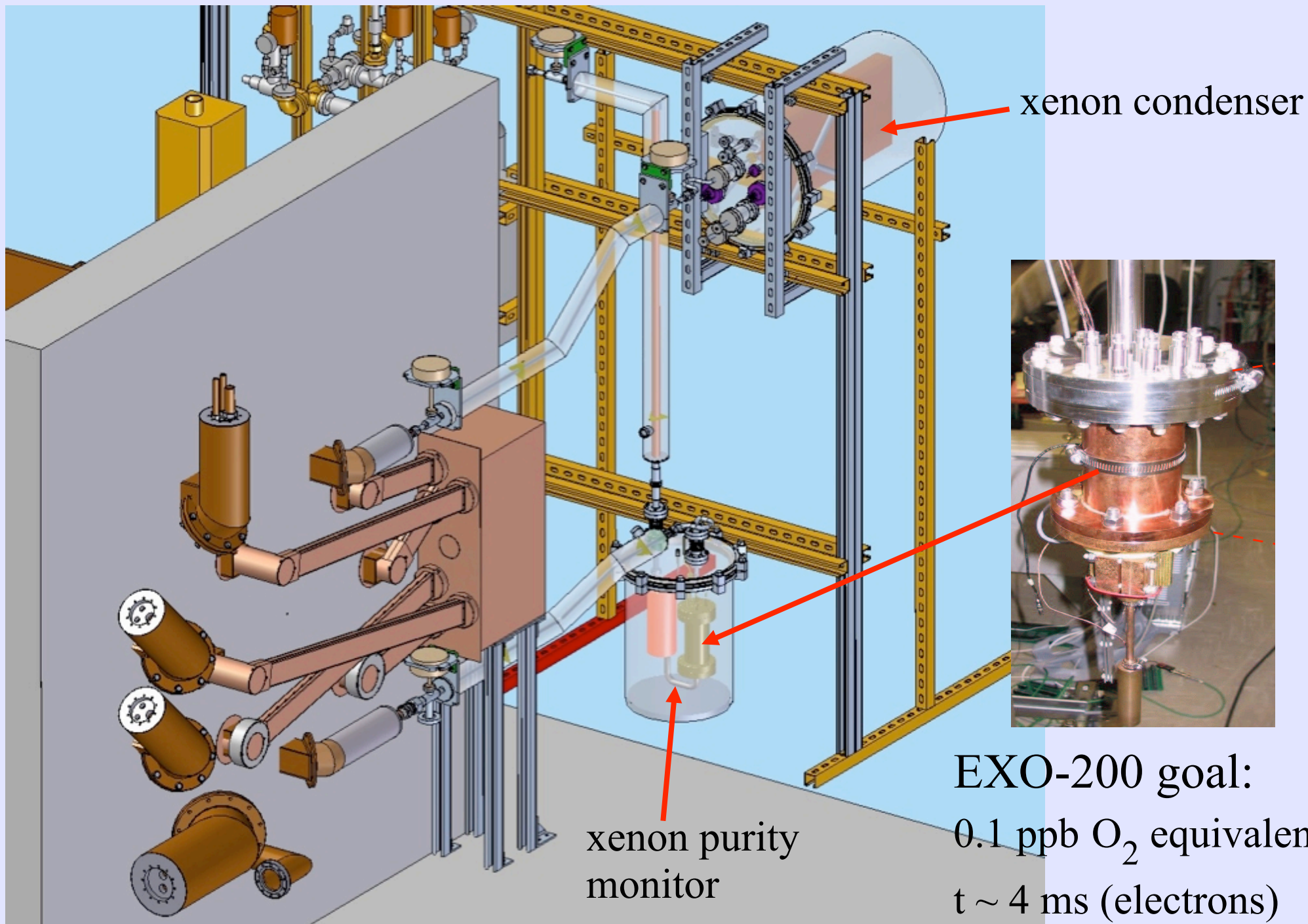
- LAAPDs stored in dry box
- First batch of 8 APDs tested and accepted last week
- 50-90 more will be received tomorrow for testing
- read out with EXO-200 electronics

LAAPD testing: data

- Setup allowed thorough testing and calibration of the EXO-200 electronics
- software developed for LAAPD DAQ is also good for EXO-200 analysis
- LAAPDs are now in production



EXO-200 installation: xenon handling system 2



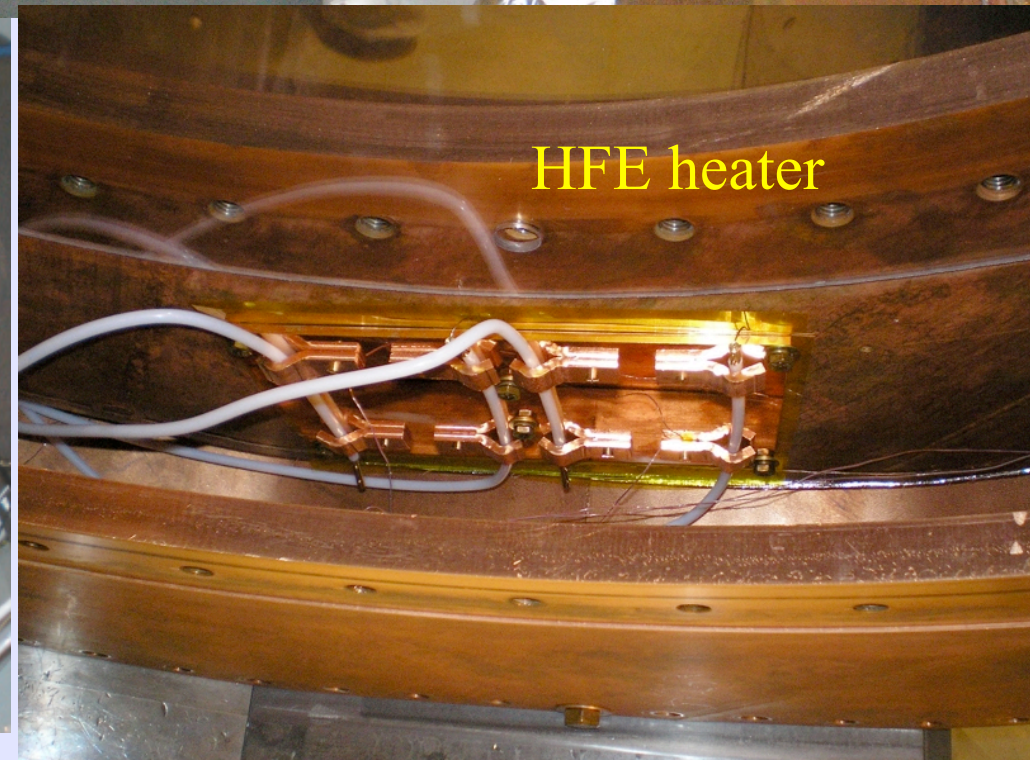
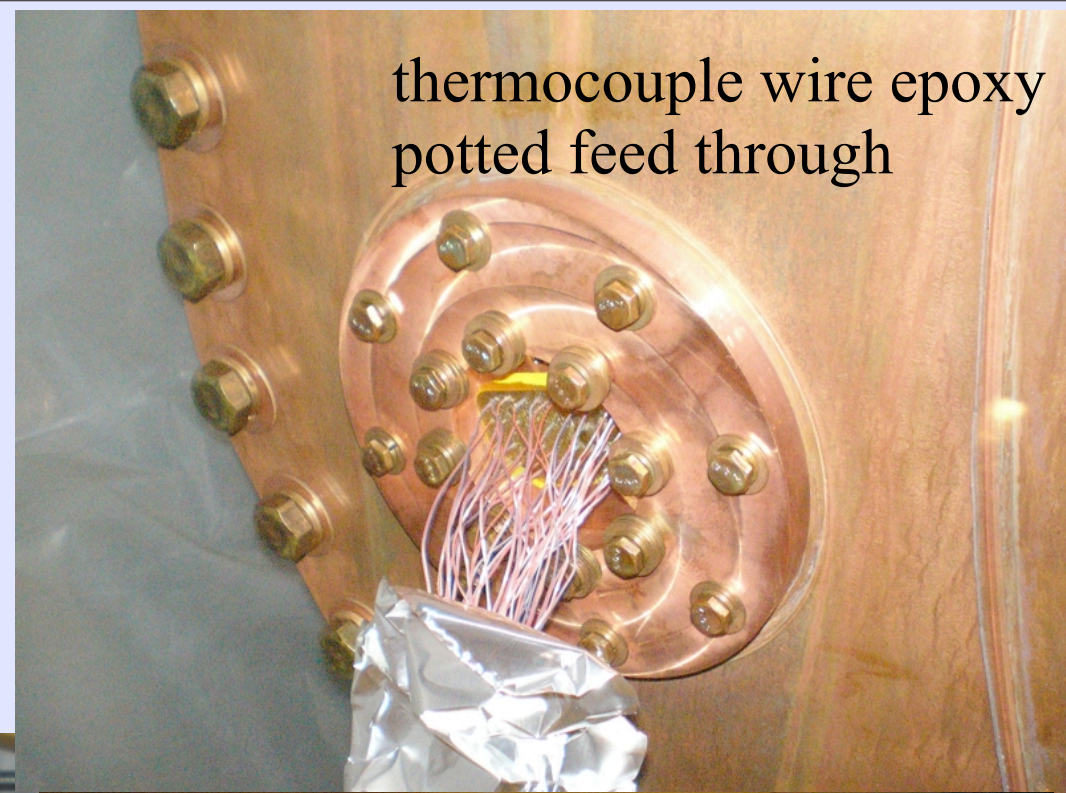
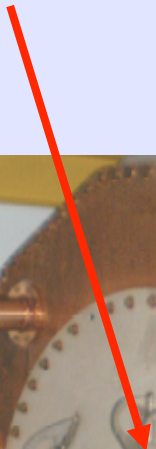
xenon condenser

xenon purity monitor

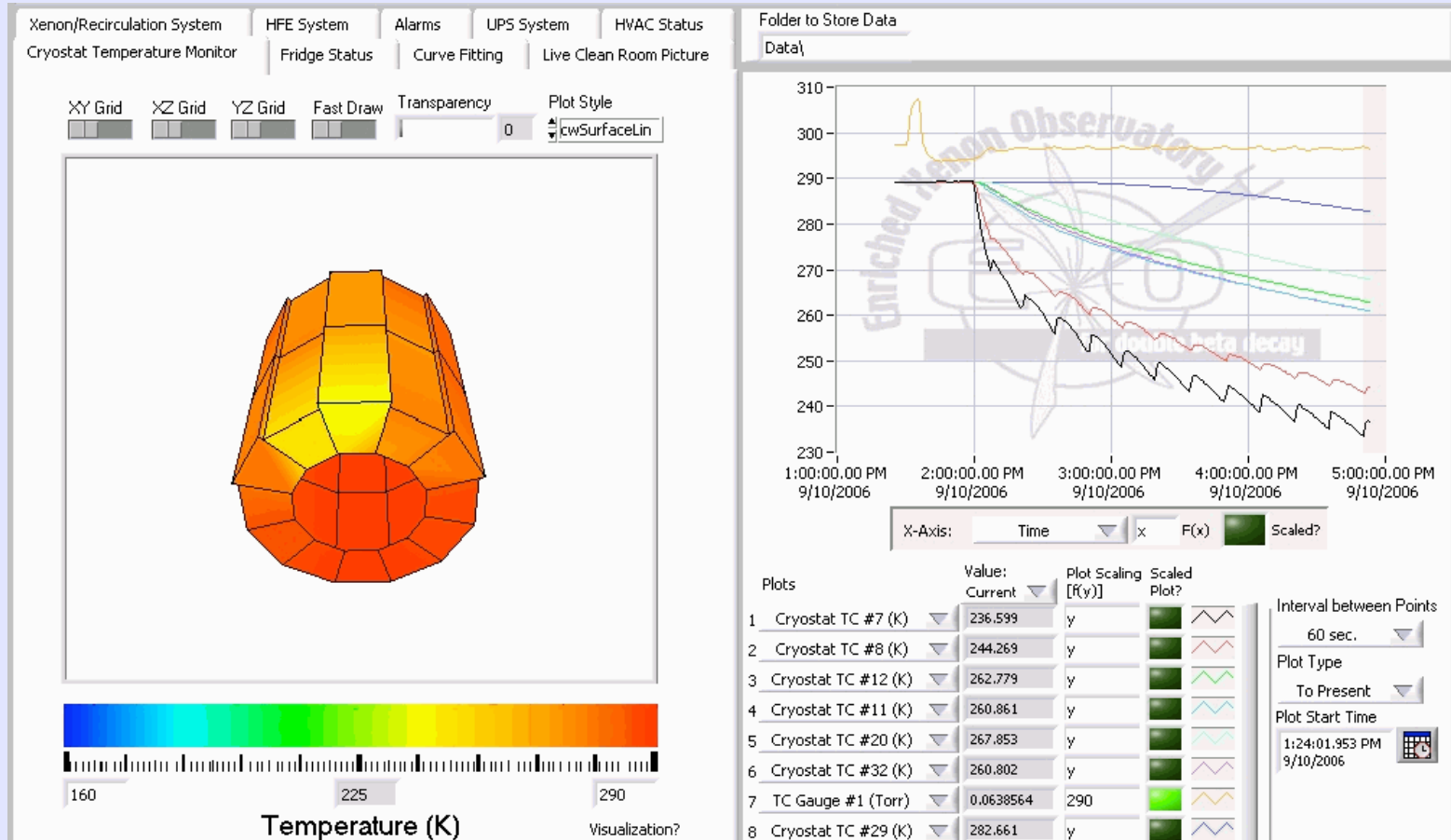
EXO-200 goal:
0.1 ppb O₂ equivalent
t ~ 4 ms (electrons)

EXO-200 installation: cryostat instrumentation

cryostat during first
cooldown and leak
checking



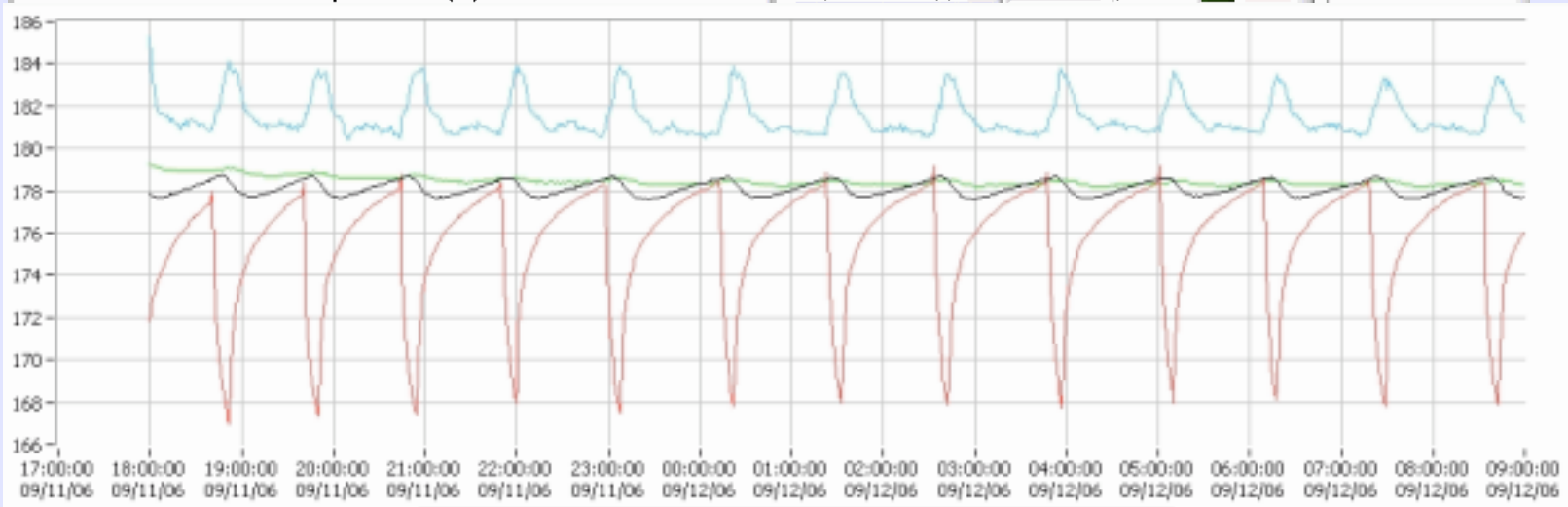
EXO-200 installation: first cooldown test



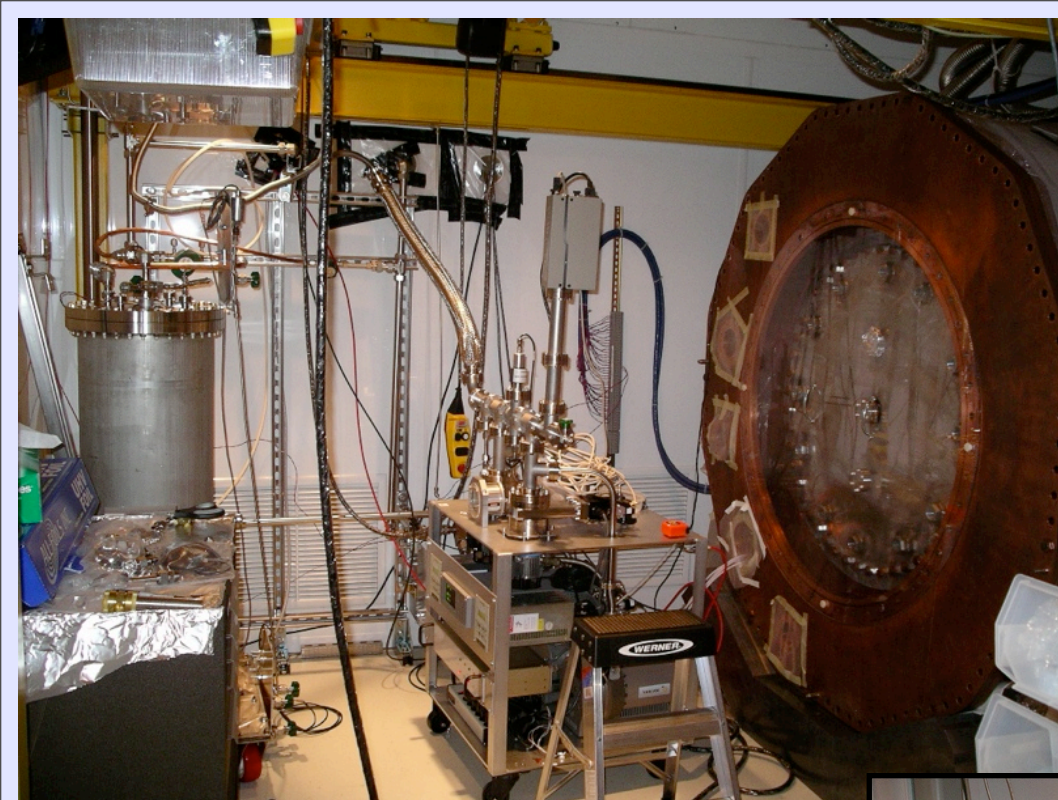
used one refrigerator

no particular problem keeping the detector at -100 C (LXe temperature)

successfully tested control system



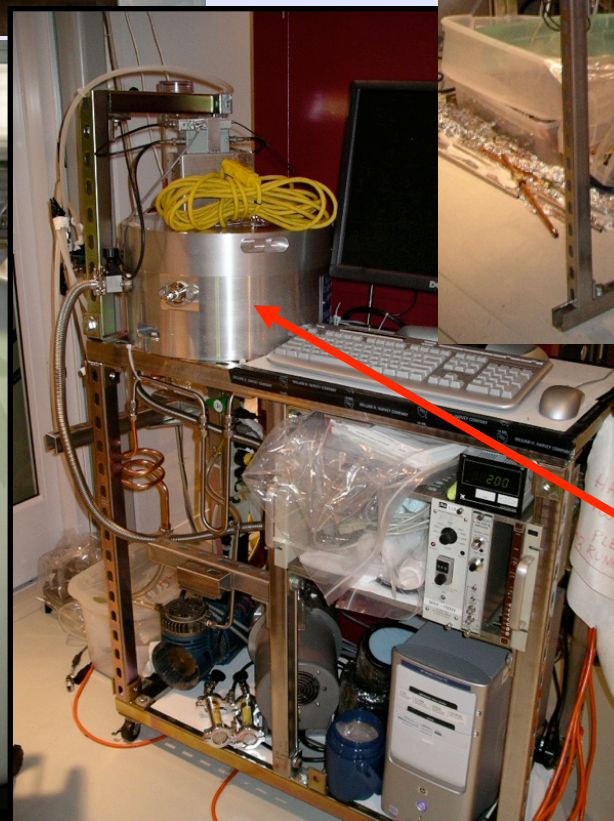
September 2006



Xe plumbing

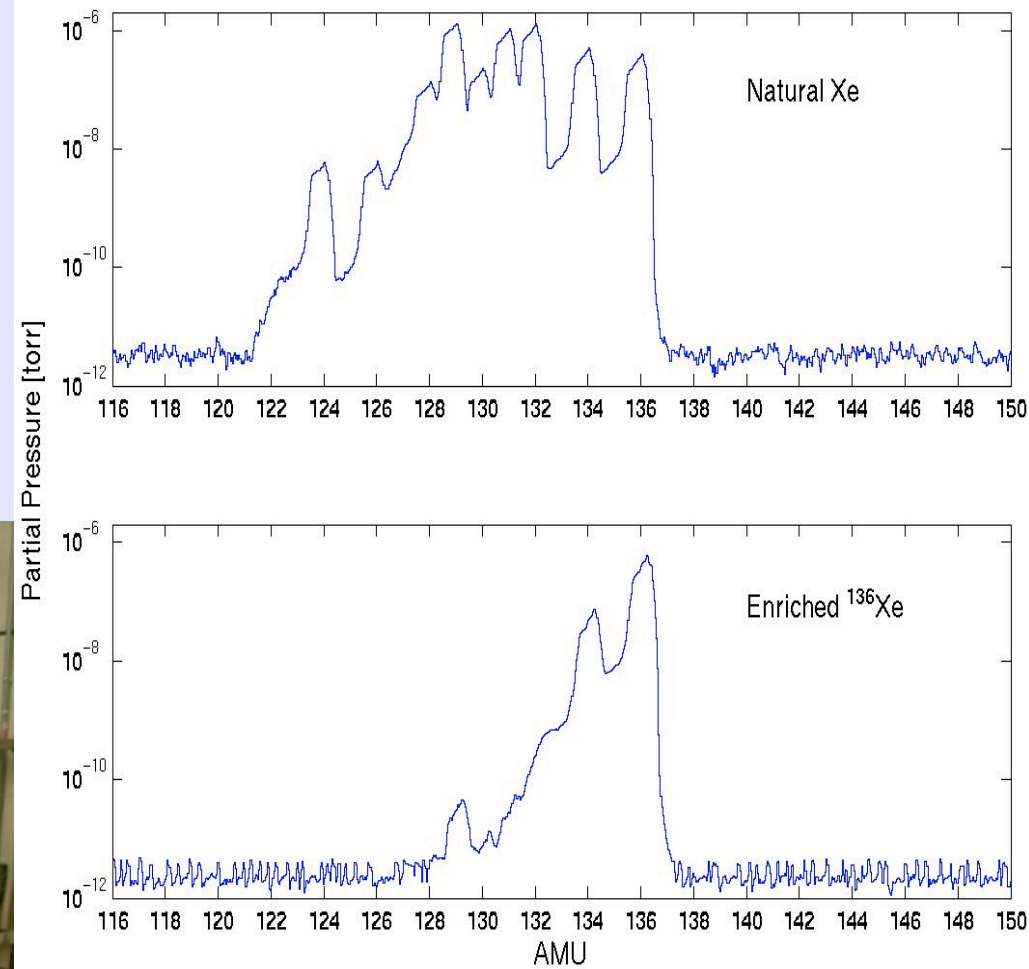
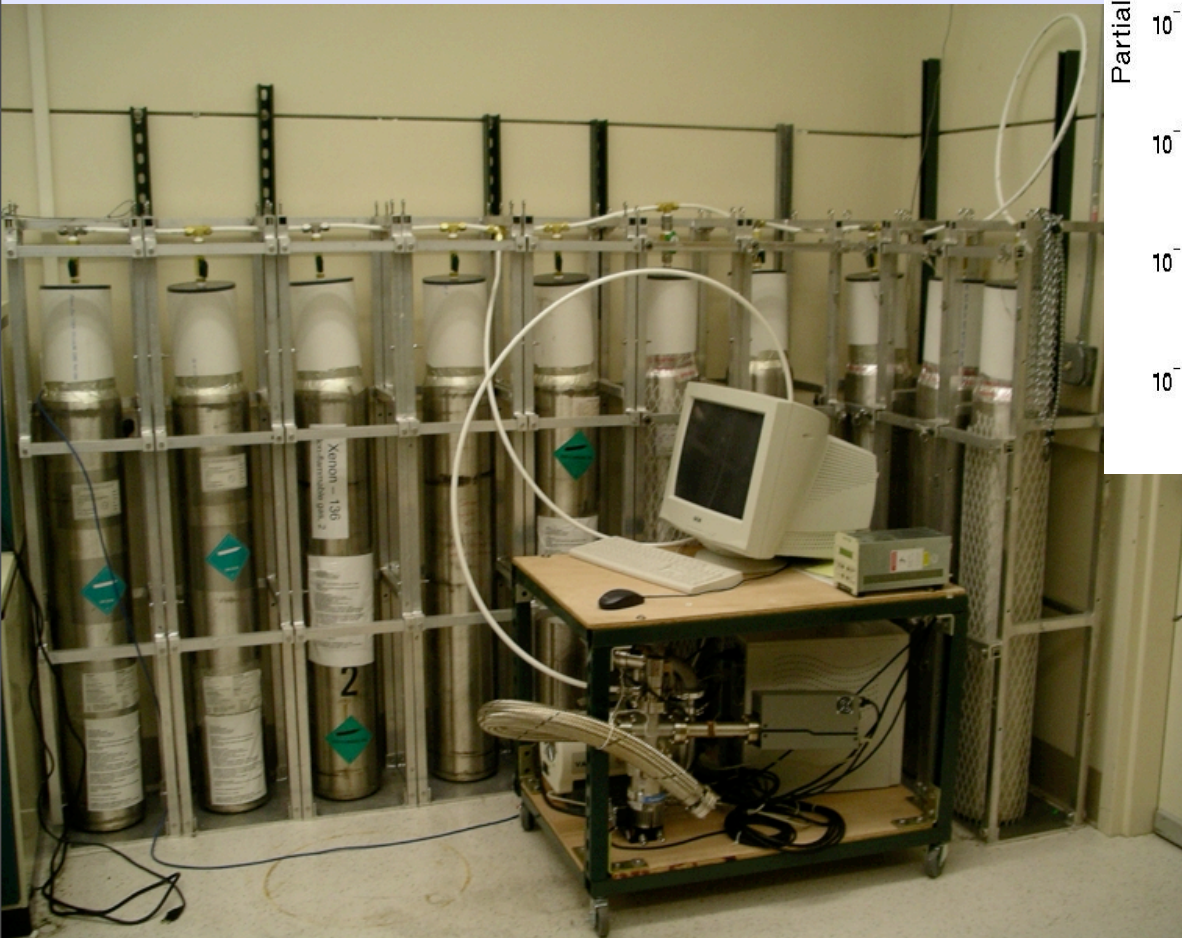


HFE plumbing



Rn emanation chamber

^{136}Xe stockpile at Stanford



200 kg of xenon enriched
to 80% in ^{136}Xe : the
most isotope possession
by any $\beta\beta$ collaboration

Trace radioactivity measurements

the material qualification campaign included:

- 1) new copper purchase for xenon vessel
- 2) etching procedures for wires and stripline cables
- 3) gold and metal analysis for APD production
- 4) silicon and phosphor bronze for TPC components
- 5) certification of cleaning and etching procedures of EXO-200 parts performed locally (acid etching in dedicated clean room at Stanford)
- 6) cryogenic epoxy
- 7) high voltage cable
- 8) APD wafers
- 9) Pb shielding selection

Goals:

- a) material selection for EXO-200
- b) most complete possible radioactivity budget to use as input of Monte Carlo simulation of backgrounds

Main γ (external) backgrounds

- γ line (2449 keV) from ^{214}Bi decay (^{238}U and ^{222}Rn)
- γ line (2615 keV) from ^{208}Tl decay (^{232}Th)
- γ line (1.4 MeV) from ^{40}K
- ^{60}Co : 1.1 + 1.3 MeV simultaneous γ 's
- other γ lines in ^{238}U and ^{232}Th chains
- other cosmogenics of Cu

Materials qualification database

- Neutron Activation Analysis (NAA) - Alabama (MIT reactor)
- ICP-MS and GD-MS - INMS (Ottawa)
- Radon emanation - Laurentian (Sudbury)
- Gamma counting - Neuchatel, Alabama
- Alpha counting - Alabama, Carleton, SLAC, Stanford

EXO Materials Testing Summary

(Status 8/31/2006)
287 entries

287 entries

Material	Information Source	MD#	K conc. [10 ⁻⁹ g/g]	Th conc. [10 ⁻¹² g/g]	U conc. [10 ⁻¹² g/g]
TPC and Internals					
SNO acrylic, batch 48, panel 09.	UA, NAA 8/26/06	59	<3.1	<16	<22
Dupont Vespel, batch SP-1 PLAQUE PGF 9713. Plaque 1. EXO production 6/22/06. Material reserved at Dupont.	UA, NAA 8/26/06	74.1	282±29	<12	<18
Dupont Vespel, batch SP-1 PLAQUE PGF 9714. Plaque 2. EXO production 6/22/06. Material reserved at Dupont.	UA, NAA 8/26/06	74.2	62±7	<25	<28
Norddeutsche Affinerie OFRP copper. Produced 6/1/2006 for EXO. Batch E263/3E1. Sample DOWN collected at DESY.	INMS (Canada) ICPMS 9/1/06	85	<55	<0.5	<0.3
	INMS				

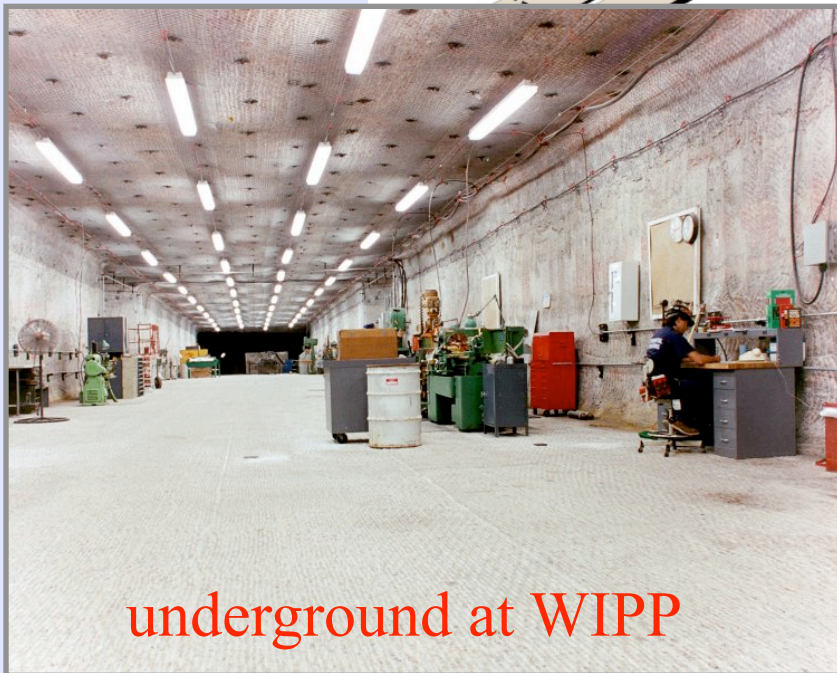
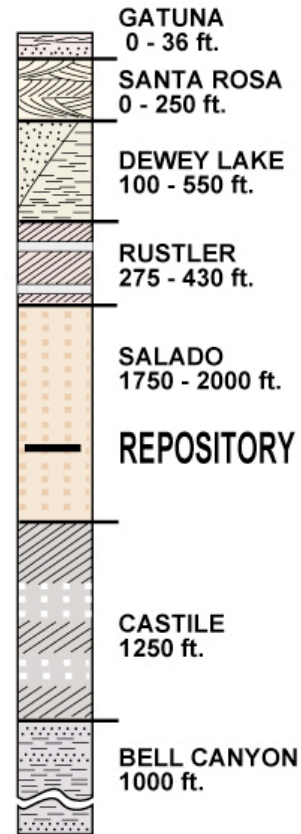
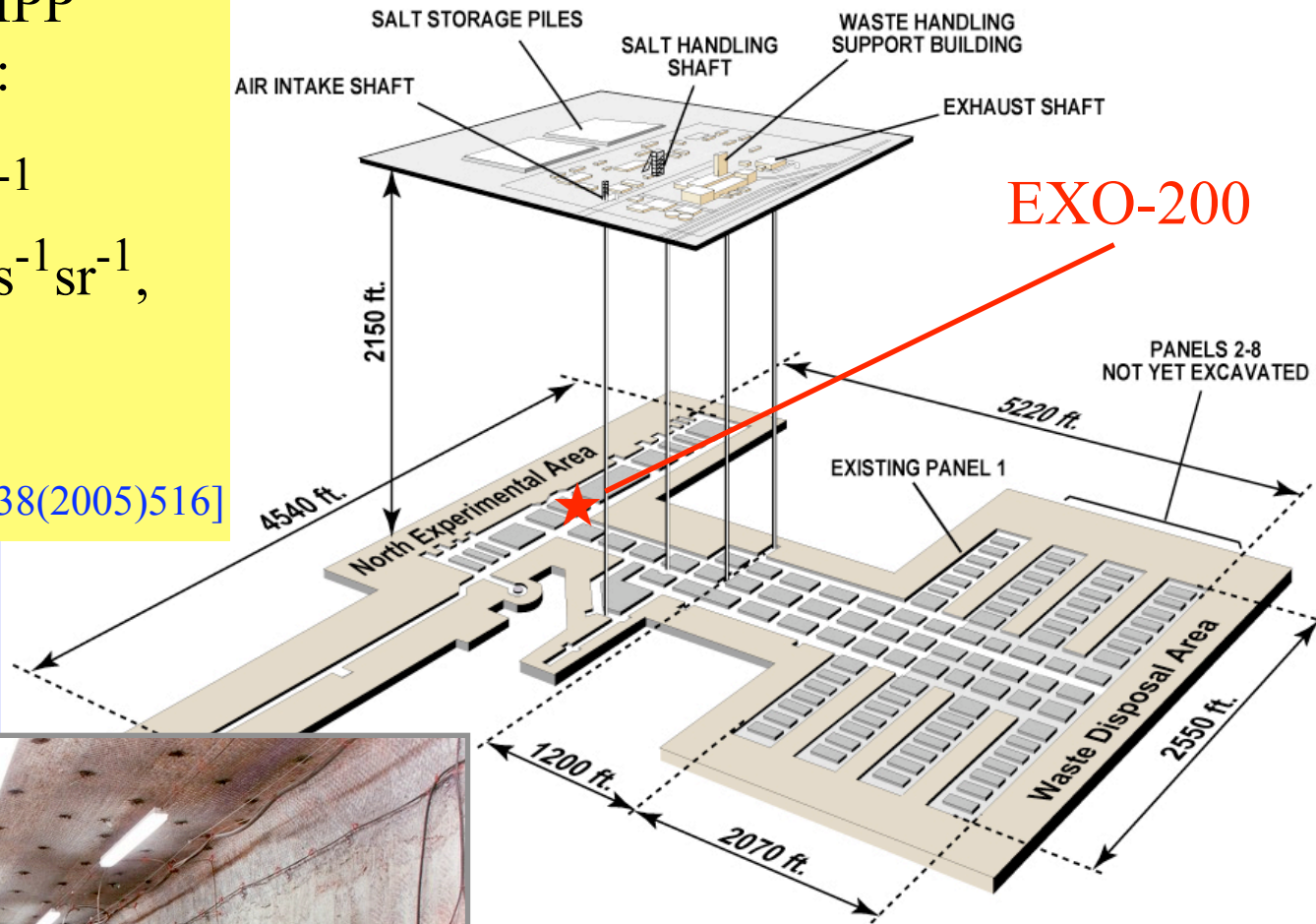
Copper shielding

- First chamber being built with leftover copper from cryostat production; this was surface shipped to Stanford and stored in ES3 (2 meters of concrete overburden).
- Cryostat copper stored in a bunker during production at SDMS, Grenoble
- 8 more tons of copper purchased and rolled in Germany; copper was stored at DESY bunker before rolling. It will be shipped to Stanford at the end of the month in a concrete shielded container (which will be used for chipping the chamber+TPC to WIPP)



WIPP Facility and Stratigraphic Sequence

muon flux at WIPP
 (~ 1700 m.w.e.):
 $4.77 \times 10^{-3} \text{ m}^{-2} \text{ s}^{-1}$
 $(3.10 \times 10^{-3} \text{ m}^{-2} \text{ s}^{-1} \text{ sr}^{-1},$
 $\sim 15 \text{ m}^{-2} \text{ h}^{-1})$
 [E.-I. Esch et al.,
 Nucl. Instr. Meth. A 538(2005)516]



underground at WIPP

EXO 200 kg prototype is being assembled and commissioned at Stanford, then the six clean rooms will be shipped to WIPP (May 2007)

EXO-200 sensitivity

Case	Mass (ton)	Eff. (%)	Run Time (yr)	$\sigma(E)/E$ @ 2.5MeV (%)	Radioactive Background (events)	$T_{1/2}^{0\nu}$ (yr, 90%CL)	Majorana mass (eV) QRPA [‡] (NSM [#])	
EXO-200	0.2	70	2	1.6*	40	6.4×10^{25}	0.18	(0.53)

* $\sigma(E)/E = 1.4\%$ obtained in EXO R&D, E.Conti et al. Phys Rev B 68 (2003) 054201

‡ QRPA: A.Staudt et al. Europhys. Lett.13 (1990) 31; Phys. Lett. B268 (1991) 312

NSM: E.Caurier et al. Phys Rev Lett 77 (1996) 1954

Improves current limits on ^{136}Xe by one order-of-magnitude

Discovery claim in Ge-76 (Phys. Lett. B 586 (2004) 198): $(0.7-4.2) \times 10^{25} \text{y}$, $\pm 3\sigma$ range

Xe-136:

$T_{1/2} = (0.58-3.5) \cdot 10^{25} \text{y}$ [Rodin et al. PRC68 (03) RQRPA] 7-43 dcs / (y 100 kg)
 $= (0.66-4.0) \cdot 10^{25} \text{y}$ [Staudt et al. EPL13 (90) QRPA]
 $= (0.48-2.9) \cdot 10^{25} \text{y}$ [Caurier et al. NPA654 (99) SM]

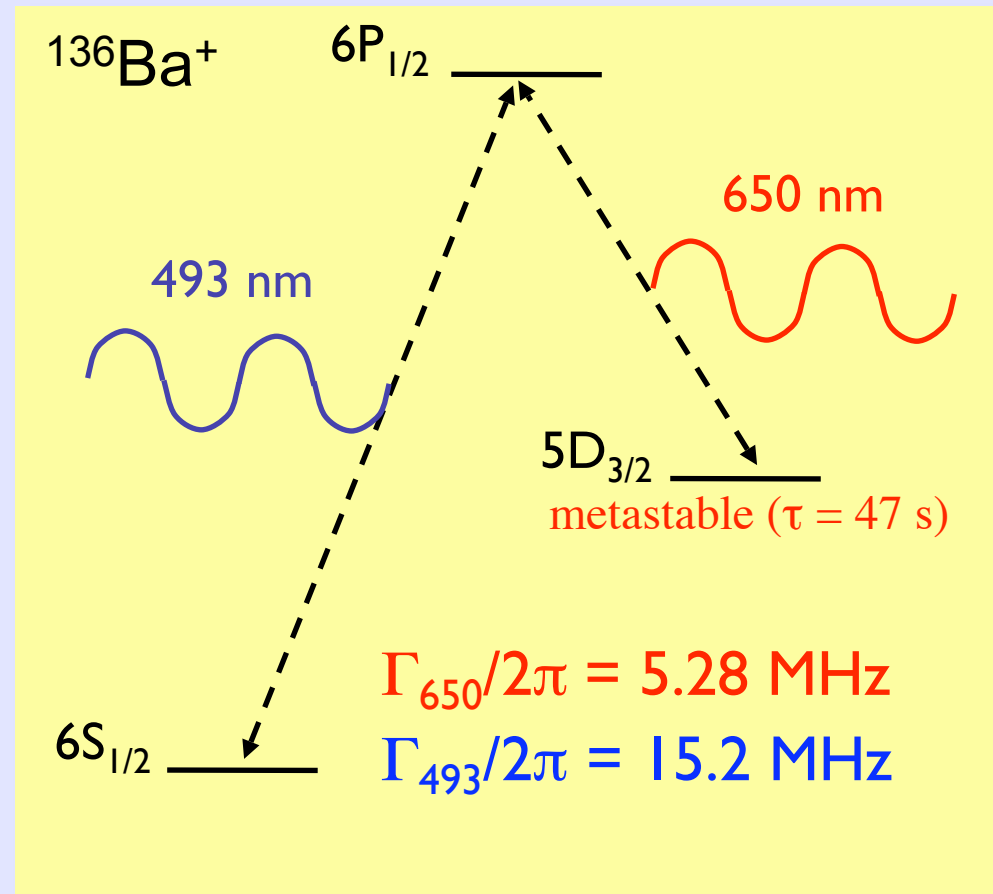
Single Ba ion detection

Daughter identified by optical spectroscopy of Ba^+ , well studied in ion traps for more than 25 years

[Neuhauser, Hohenstatt, Toshek, Dehmelt, Phys. Rev. A 22 (1980) 1137]

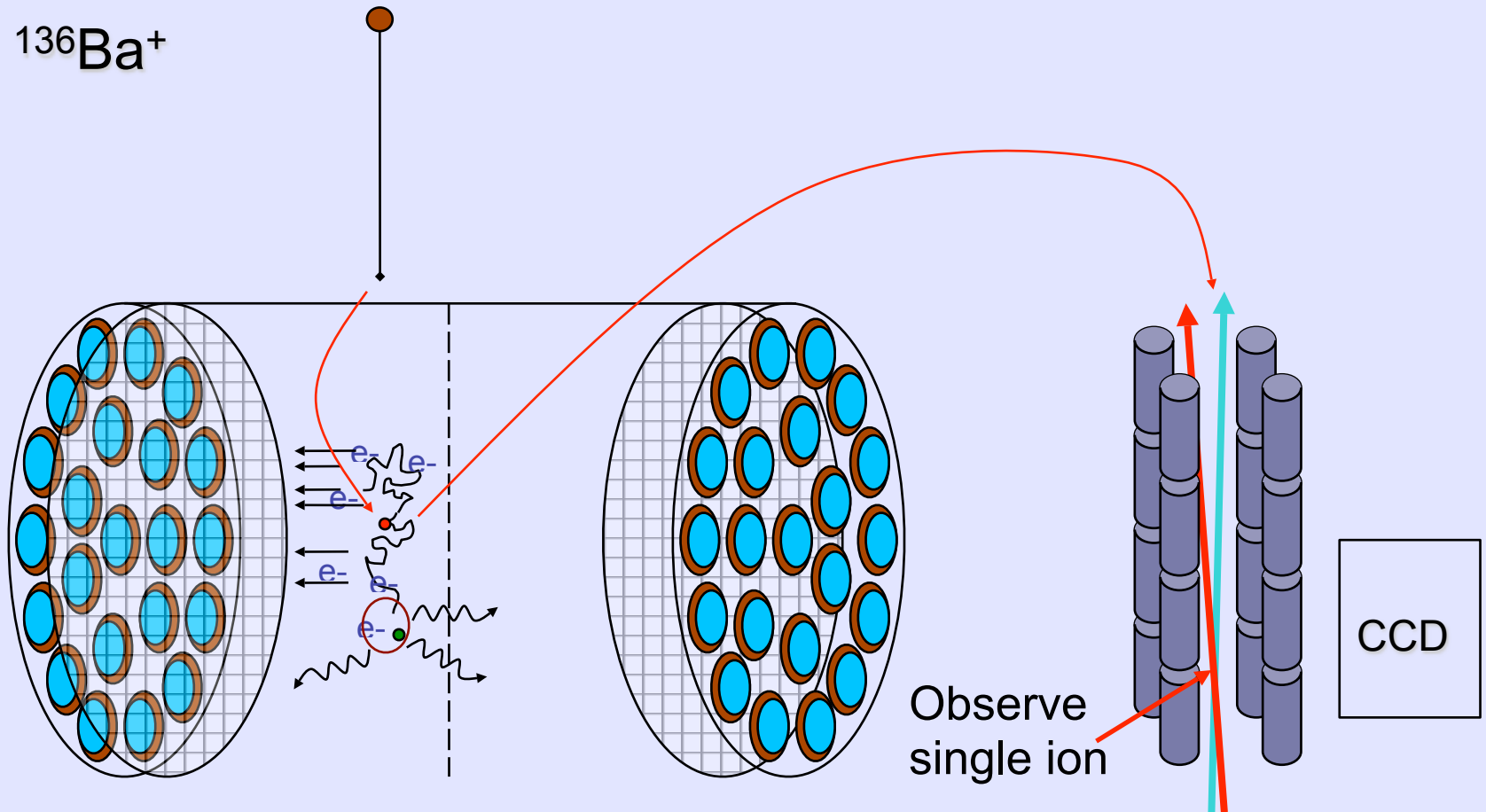
- very specific signature
- Cycling 493/650 nm transitions gives a fluorescence rate of $\sim 10^7$ Hz (in vacuum)

bright!



Ba identification strategies

- grab from Xe bath, release in a trap, and identify it



- in situ detection:
 - in LXe: Bill Fairbank at Colorado State University
 - in GXe: David Sinclair at Carleton University and SNOLab

Ba tagging requirements

We need to be able to:

- Identify a single $^{136}\text{Ba}^+$ ion
- Inject an ion into a trap from an external source
- Trap with high overall efficiency
- Trap in the presence of some Xe

Single Ba ion trapping

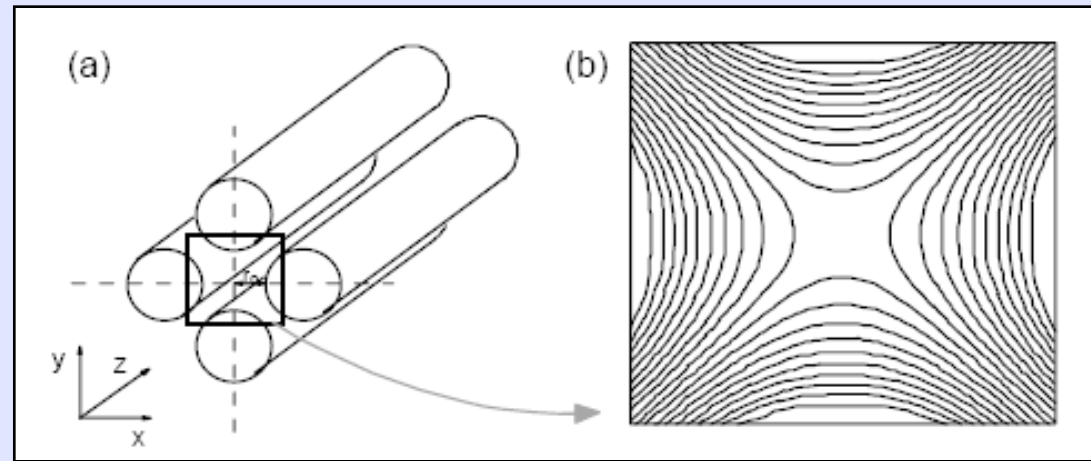
Build an AC quadrupole potential:

$$\Phi = \frac{\varphi_0}{2} \left(1 + \frac{x^2 - y^2}{r_0^2} \right)$$

$$\varphi_0 = U_{DC} + V_{RF} \cos \Omega t$$

longitudinal
trapping

radial trapping



write: $\vec{F} = q\vec{E} = m\vec{a}$

$$m \begin{pmatrix} \ddot{x} \\ \ddot{y} \end{pmatrix} = \begin{pmatrix} -\frac{e\varphi_0}{r_0^2} x \\ +\frac{e\varphi_0}{r_0^2} y \end{pmatrix}$$

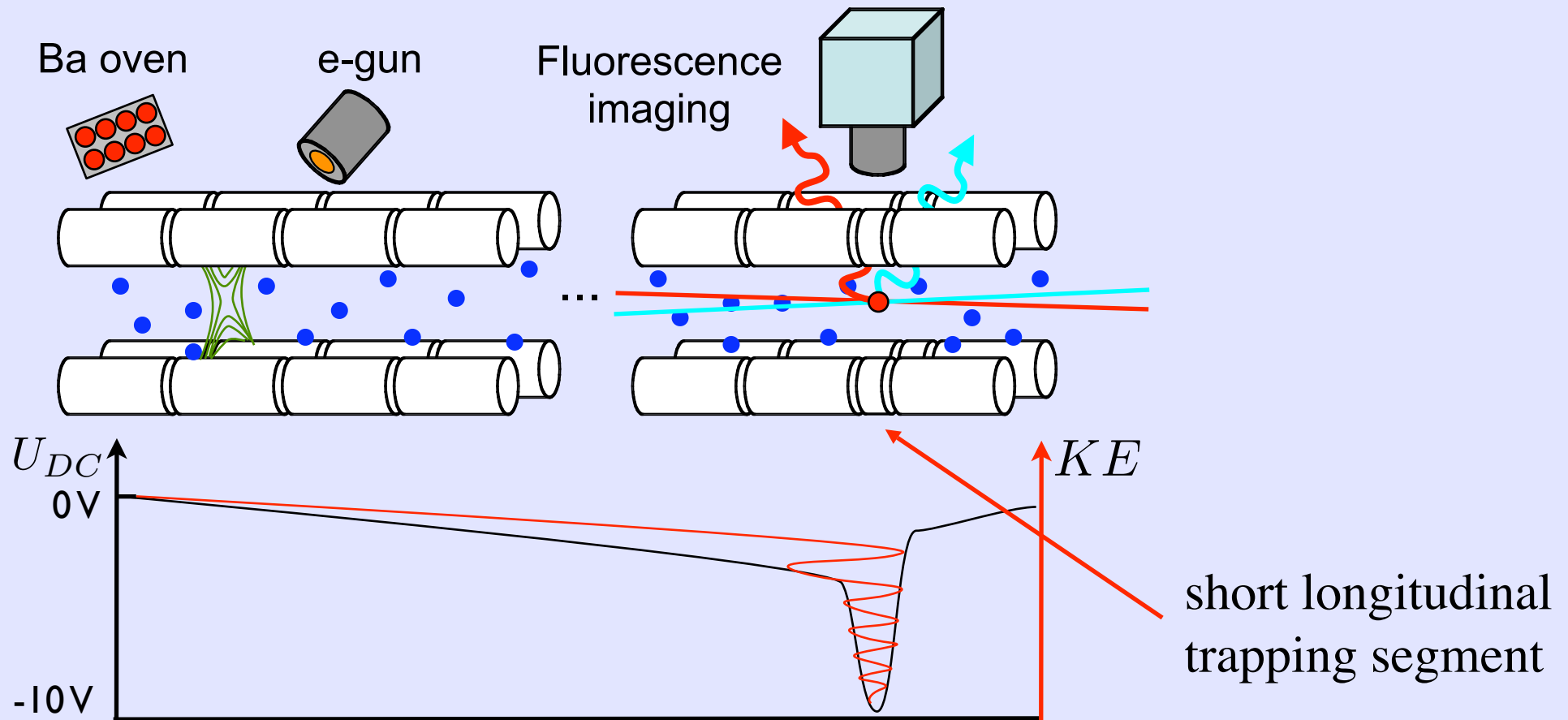
trap parameters:

$$V_{RF} = 150V_{pk}, \quad f = 1.1\text{MHz}$$

$$U_{DC} = 10V$$

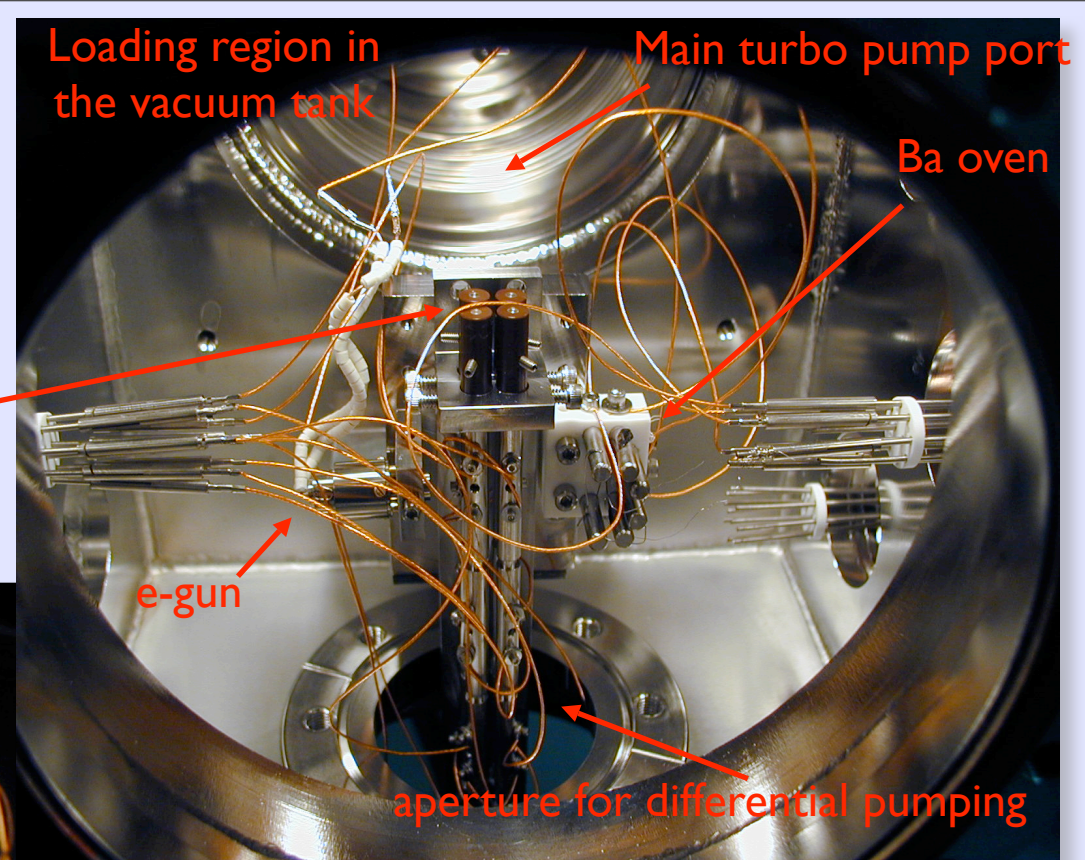
Single Ba ion trapping

Multiply by 16, and add a buffer gas to cool down the ions injected at one end of the trap into a DC minimum

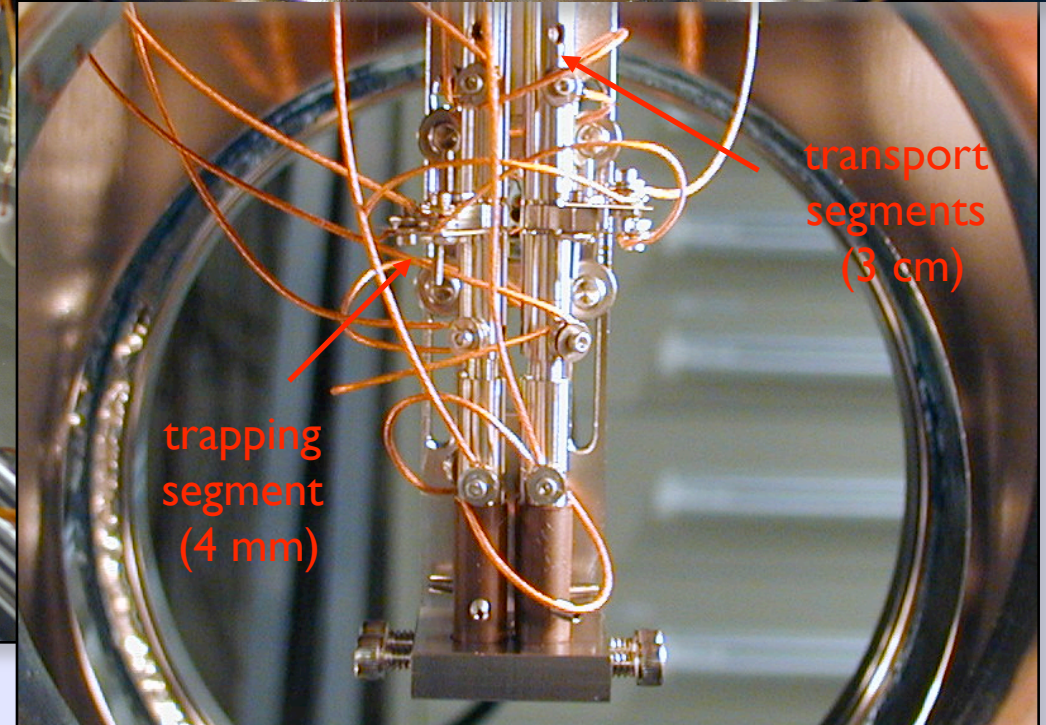


Linear ion trap at Stanford

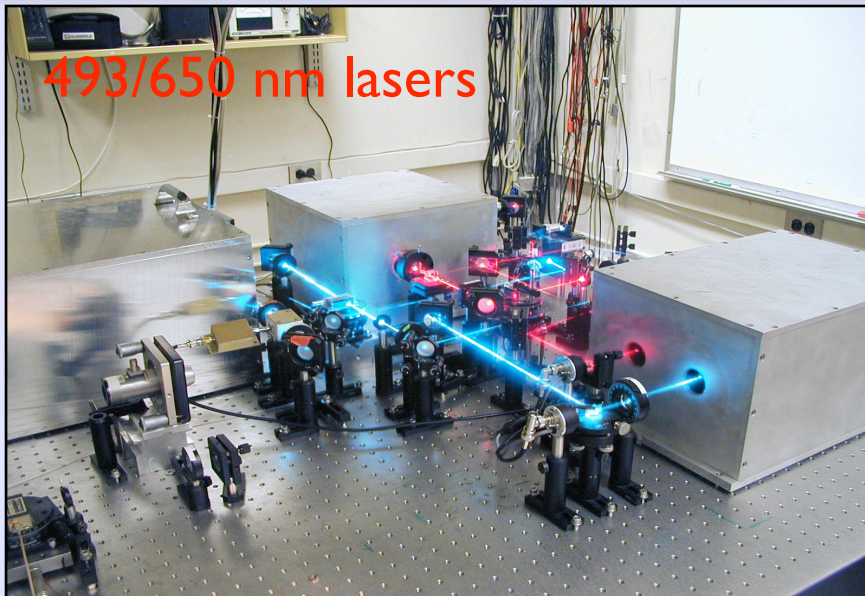
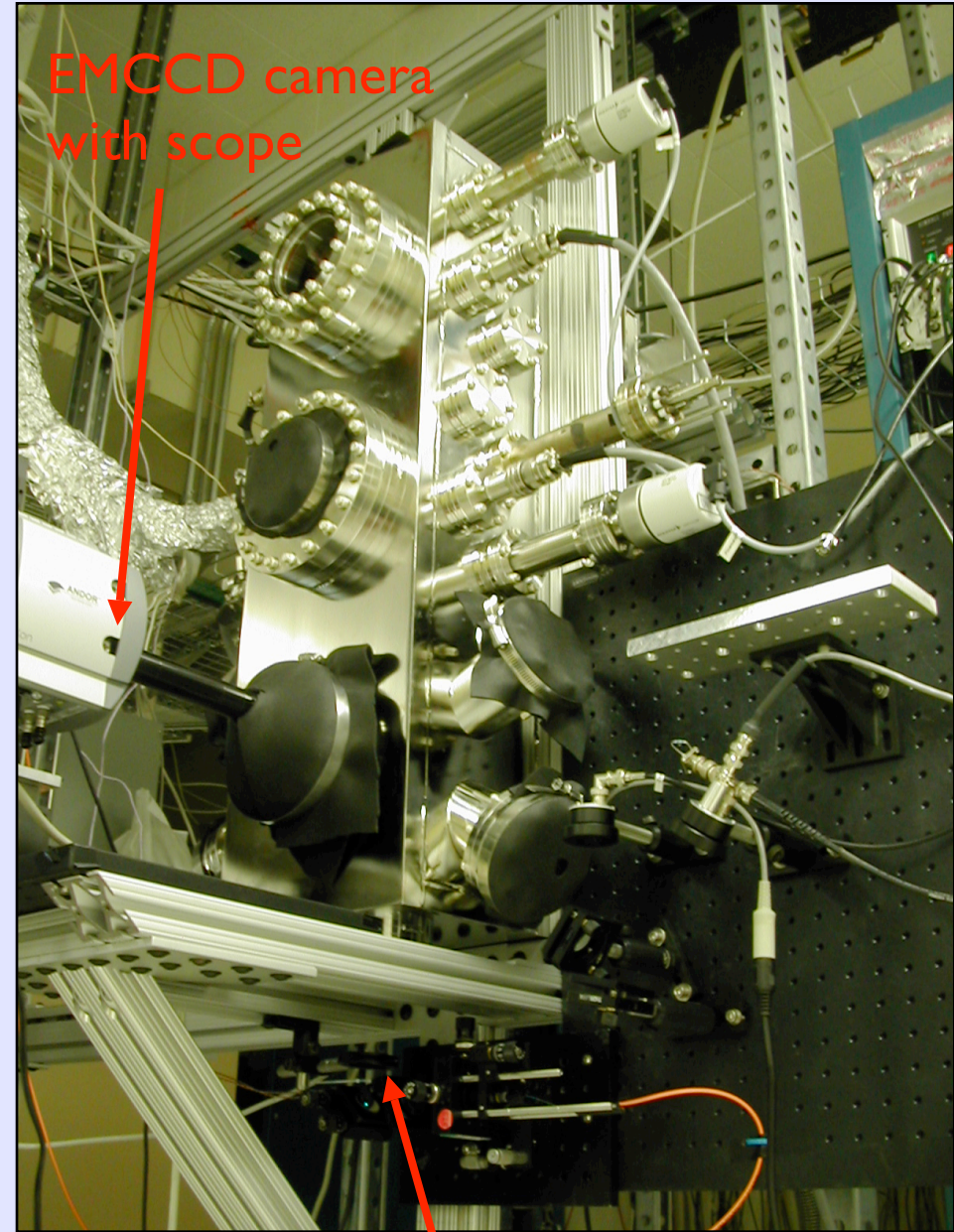
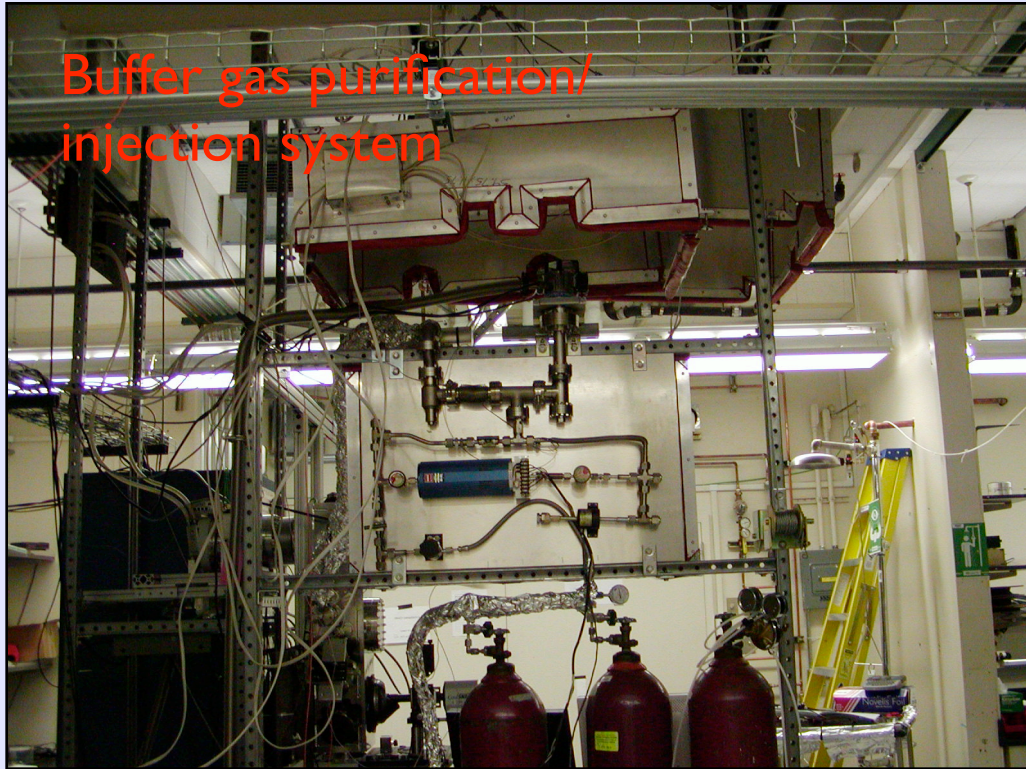
Tip loading access



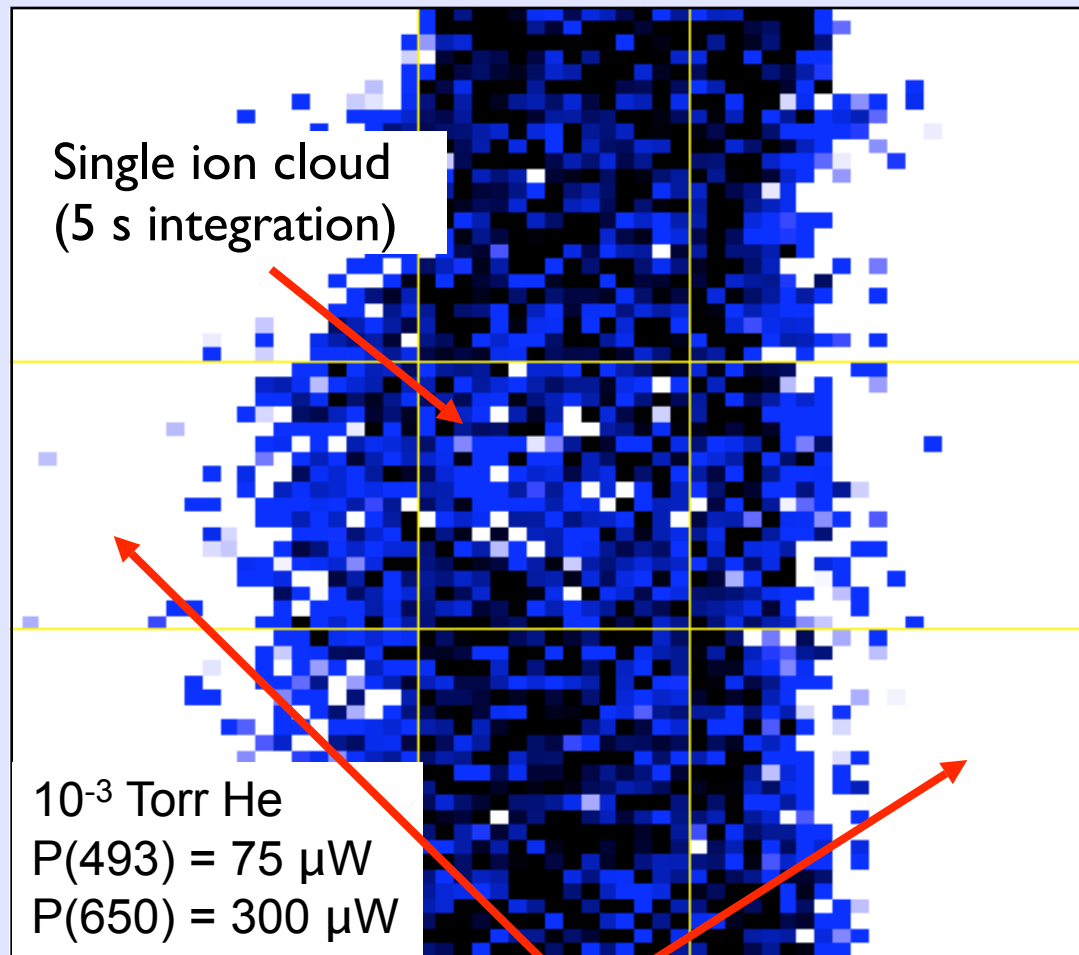
Electrode structure
being prepared



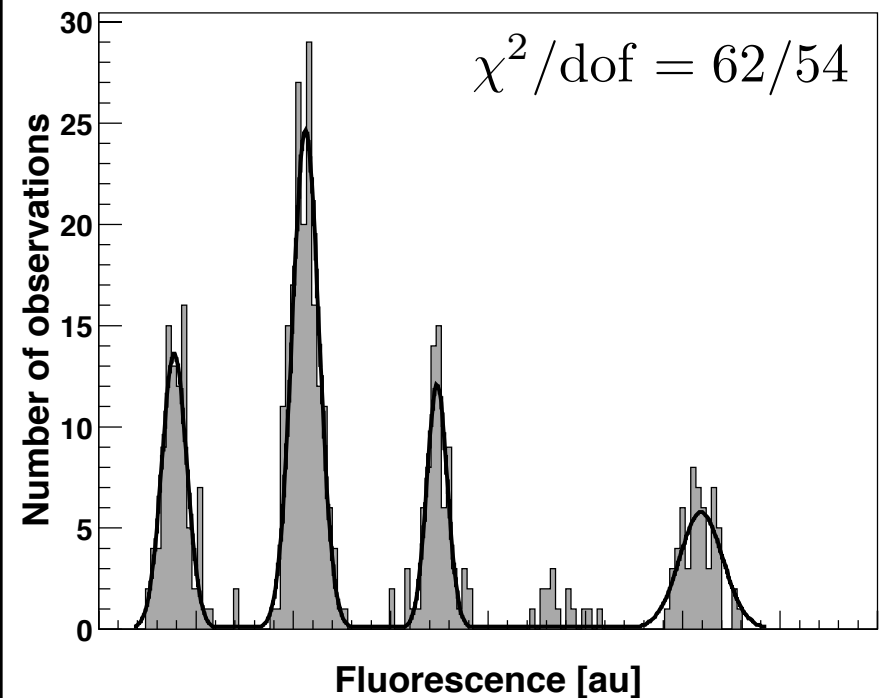
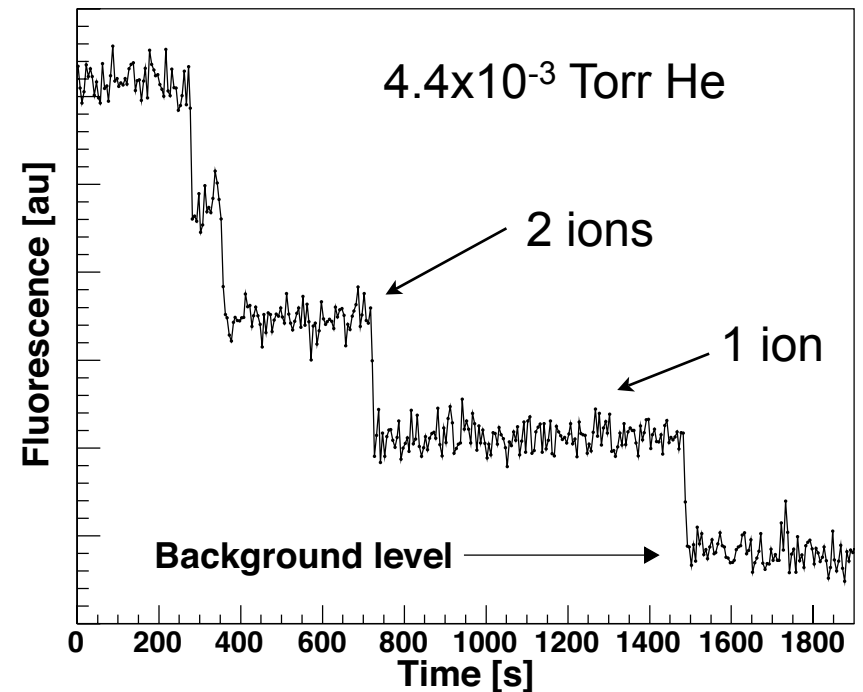
Linear ion trap at Stanford 2



Single ion detection in buffer gas



Electrodes glowing from
scattered laser light

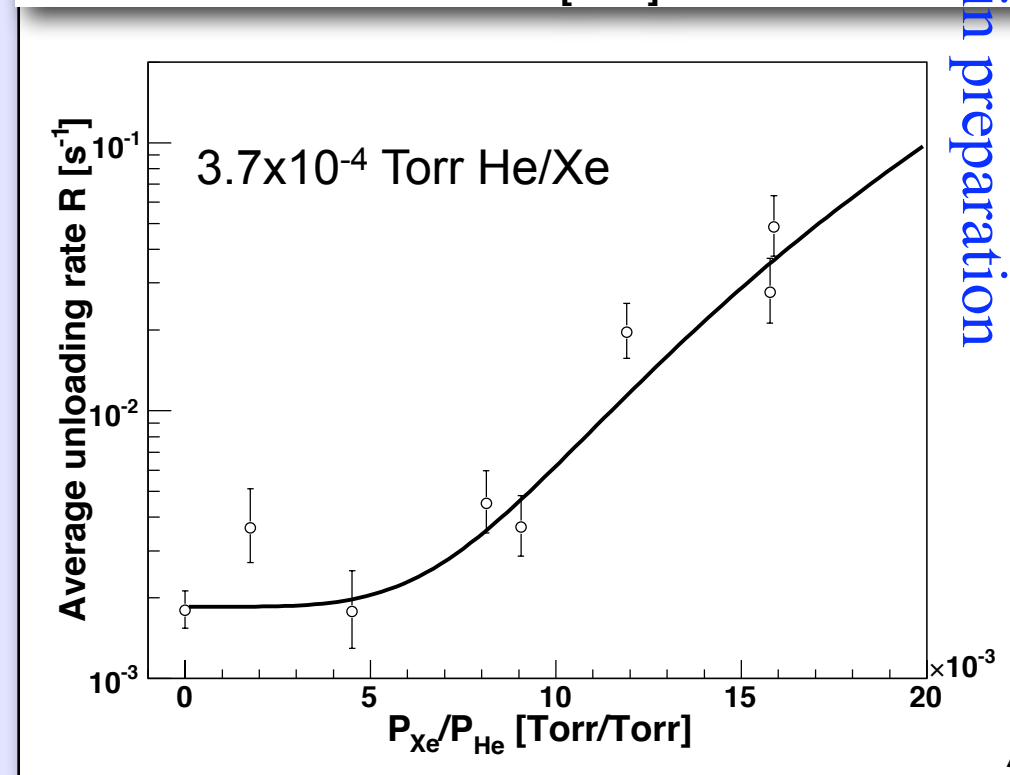
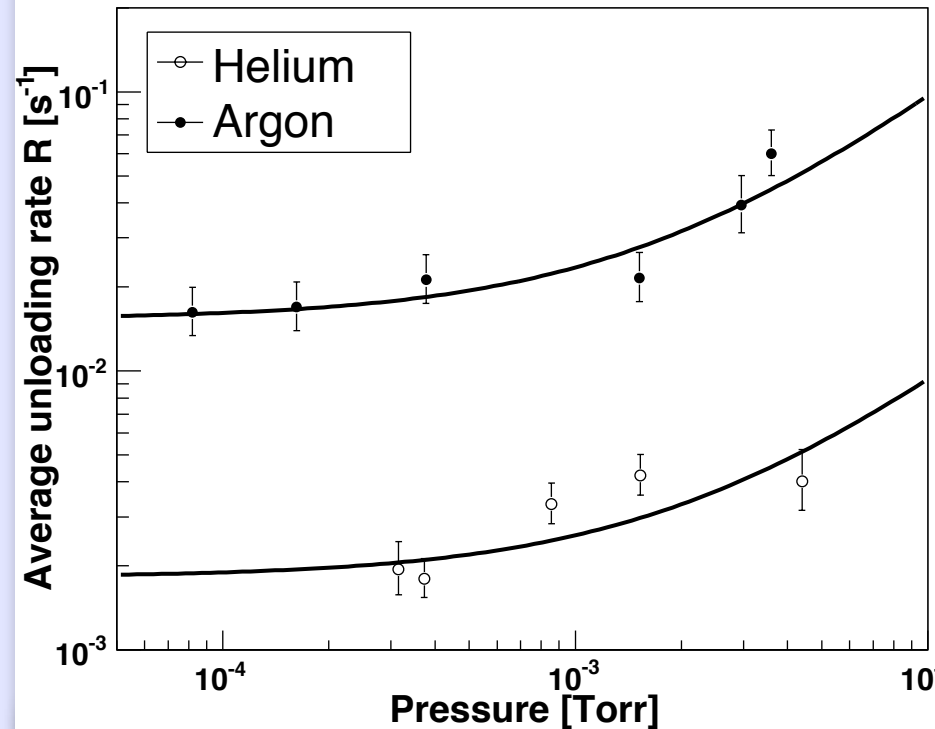


Unpublished, paper in preparation

Single ion lifetime in buffer gas

single Ba^+ ions live:
 ~ 5 min in He
 ~ 1 min in Ar

should operate at $< 5 \times 10^{-3}$
Xe/He concentrations
(Xe can unload Ba^+ with
few collisions)



Unpublished, paper in preparation

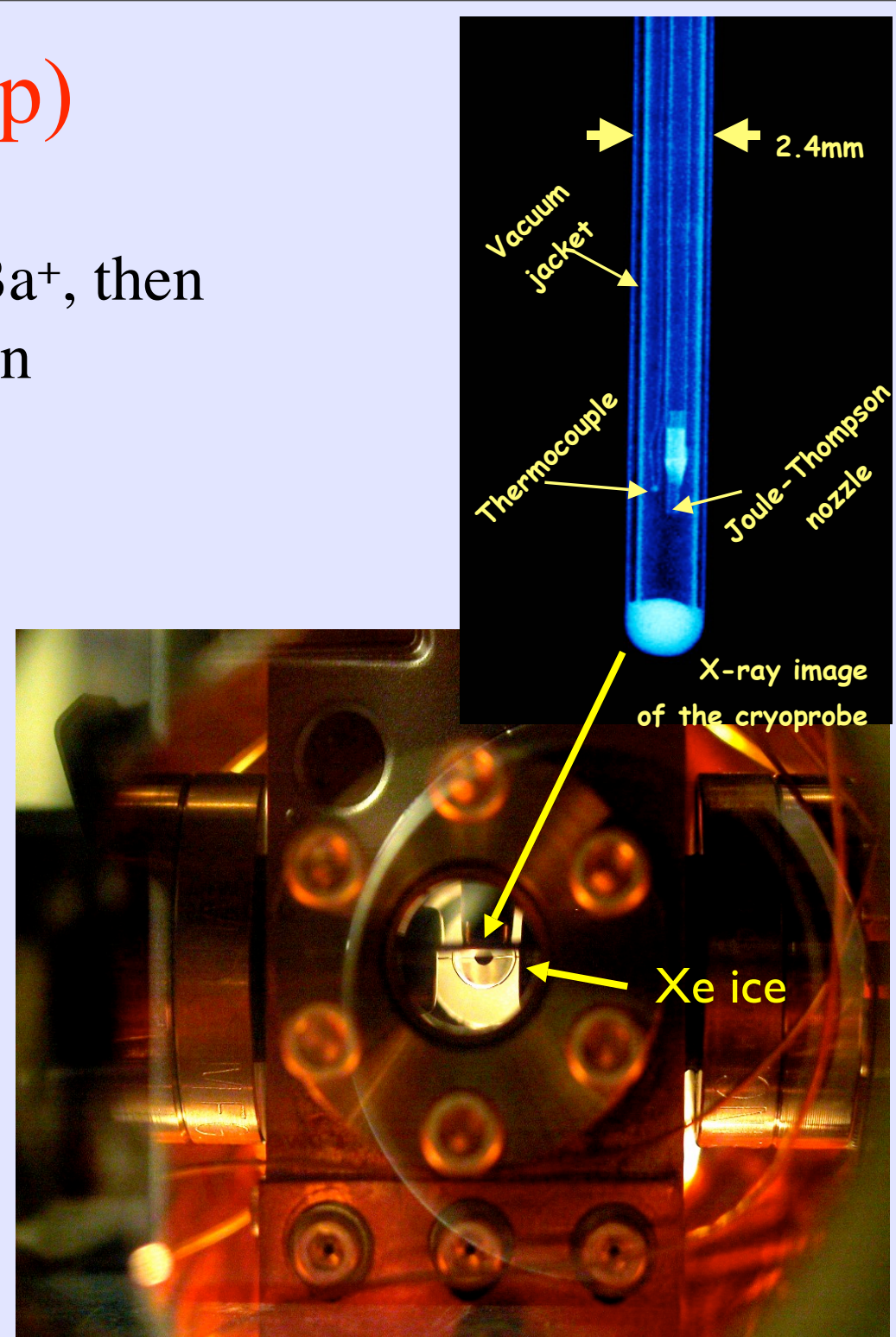
Ba ion grabbing tips

- Cold tip: freeze Ba^+ in Xe ice, then sublime and release inside trap
- Resonant ionization tip: attract Ba^+ to metalized fiber optic tip, then desorb and ionize with laser pulses inside trap
- Hot tip: attract Ba^+ to metal tip, then desorb inside trap by heating
- Field emission tip: attract Ba^+ to extremely sharp metal tip, then desorb inside trap

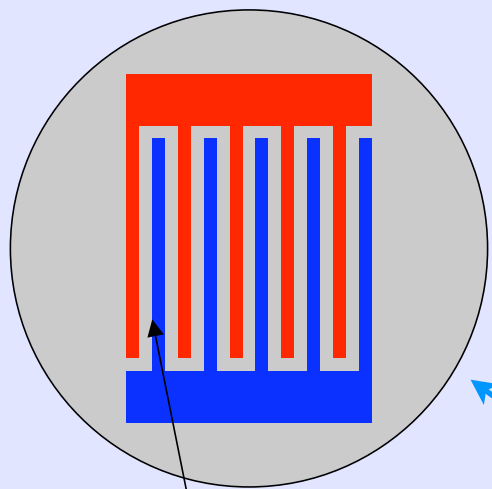
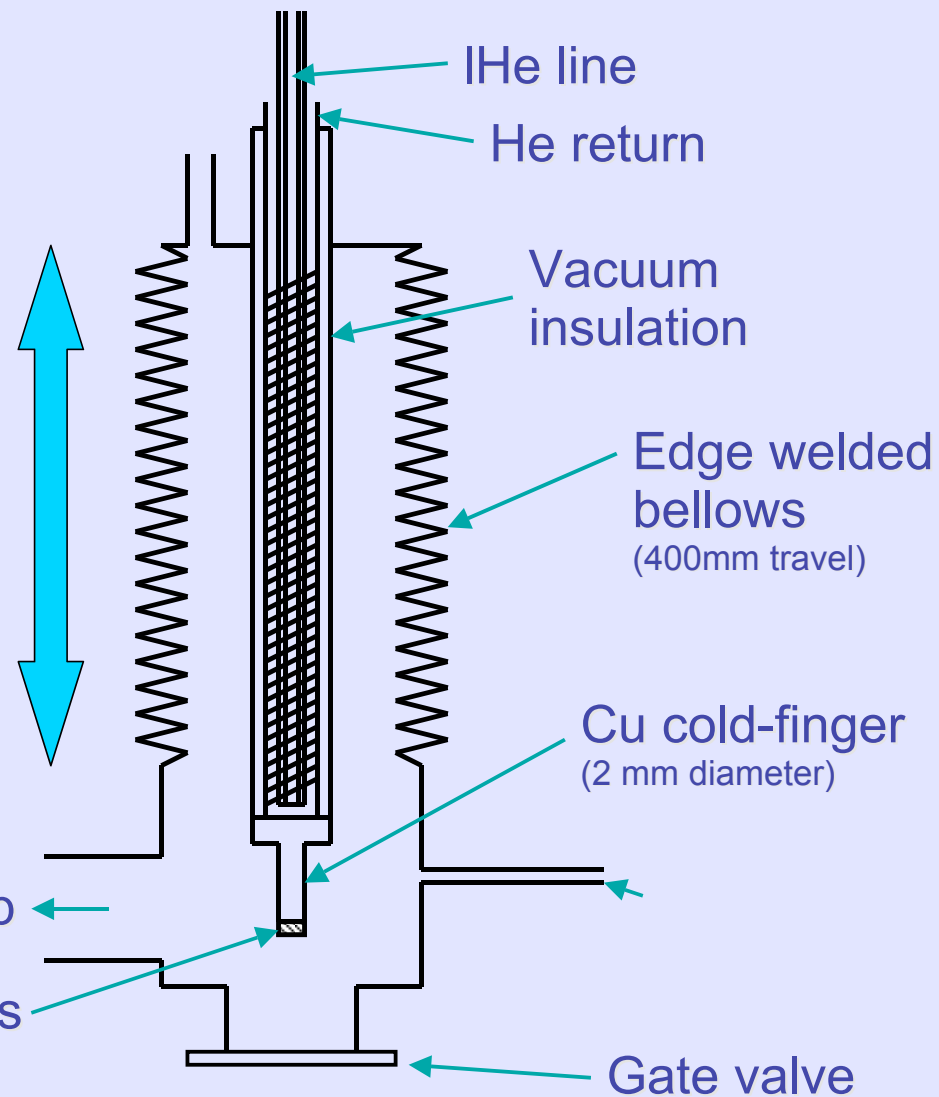
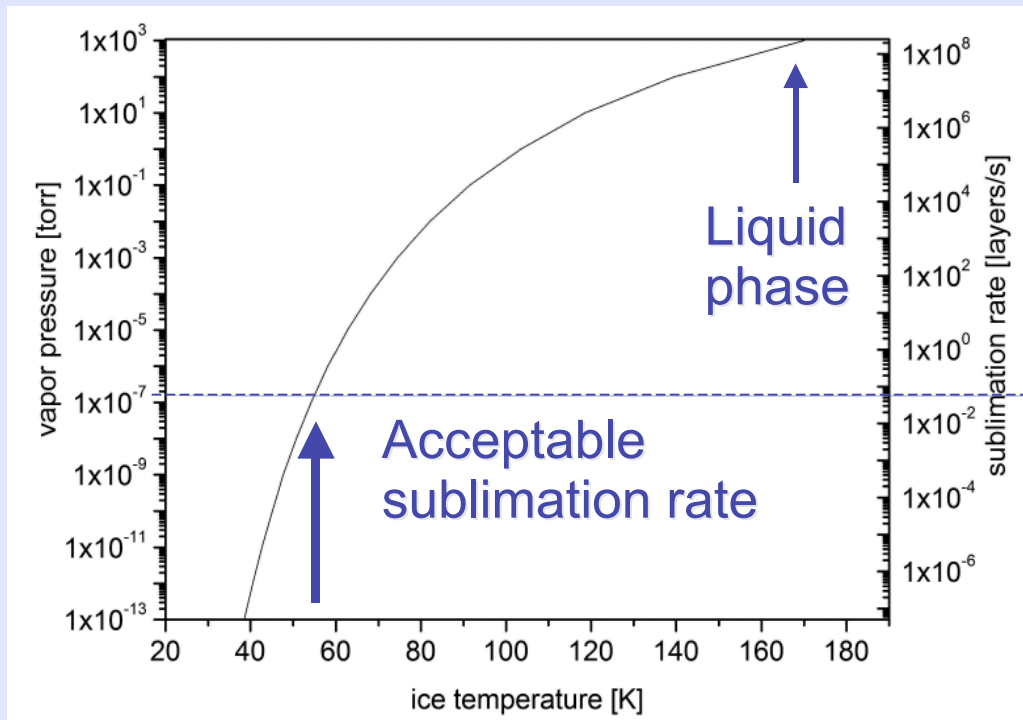
Cold ion tip (cryo-tip)

Concept: freeze Xe ice around Ba^+ , then release inside trap by sublimation

- demonstrated ability to capture and release Th^+ and Ra^+ ions
- need to control ice formation and melting to preserve the ion and minimize Xe inside the trap (thin Xe ice layer)



Capacitive cryo-tip



'fingers' for dielectric ice-thickness measurement:

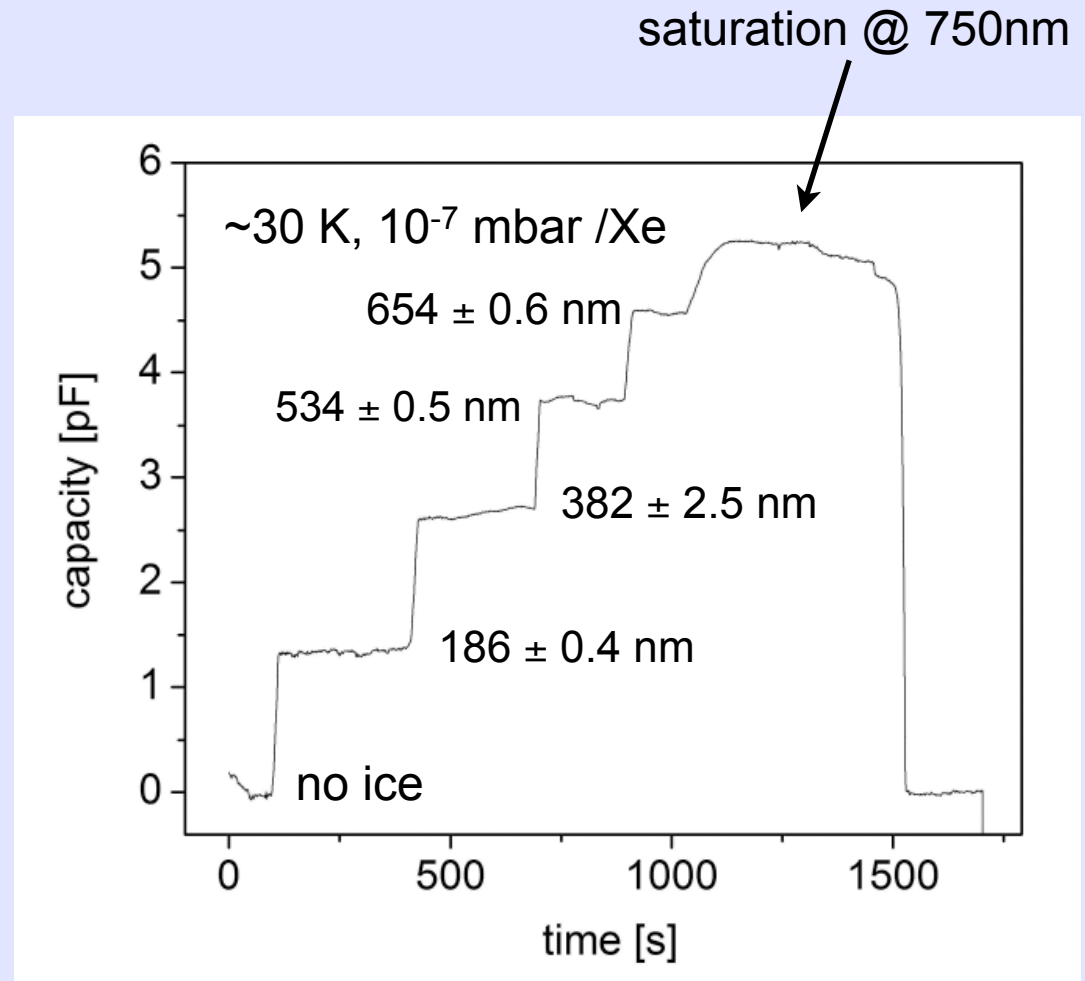
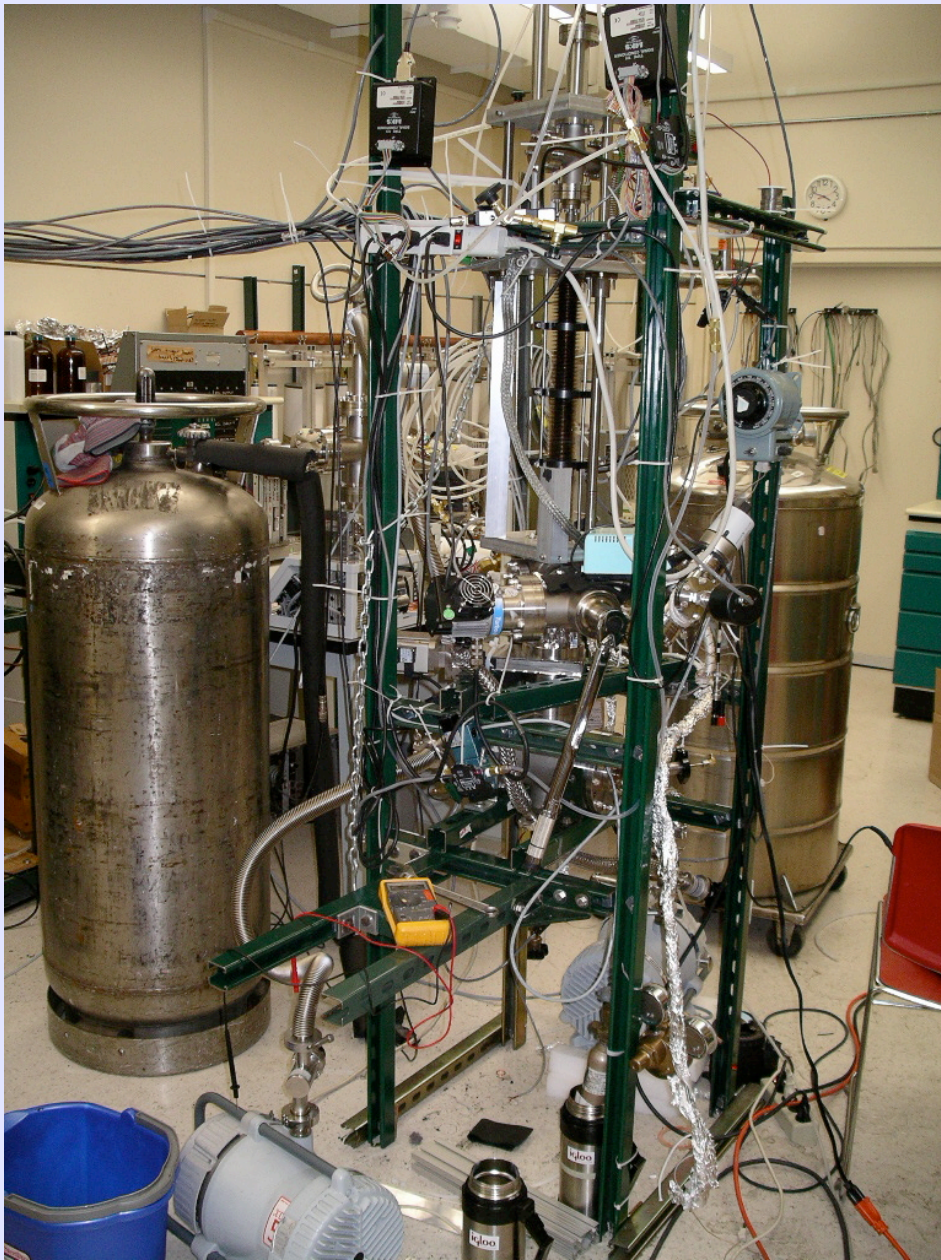
$$\epsilon_r(\text{Xe, liquid}) = 1.88$$

$$\epsilon_r(\text{Xe, solid}) = 2.25$$

Turbo pump
Ice thickness sensor

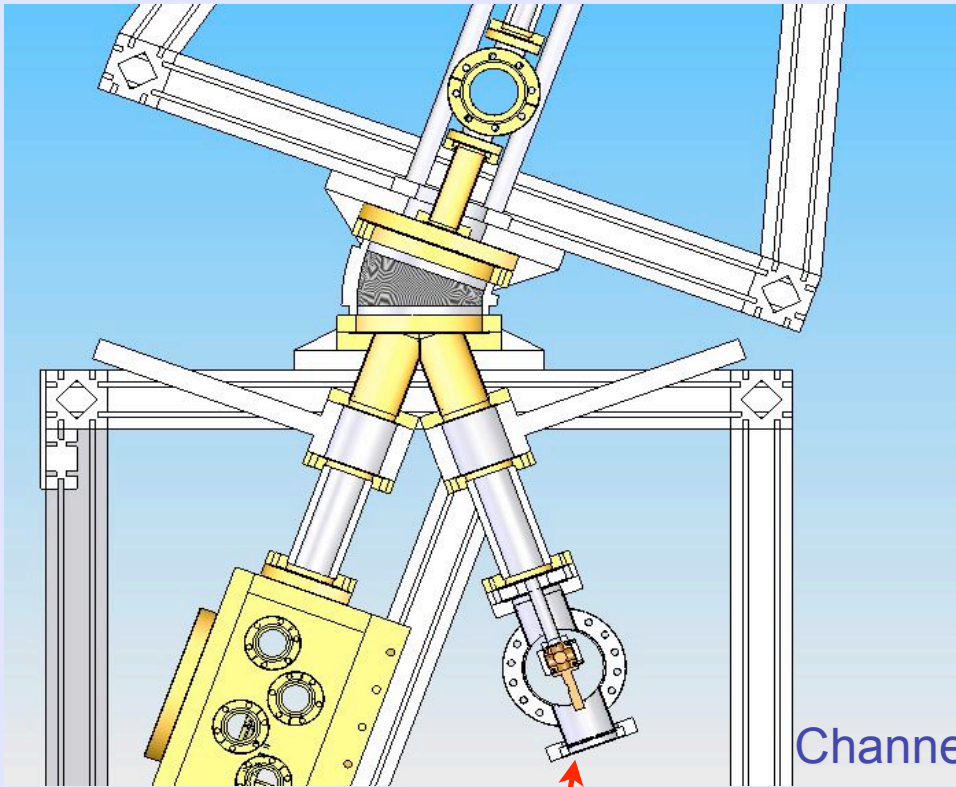
Gate valve

Thin Xe ice with capacitive cryo-tip



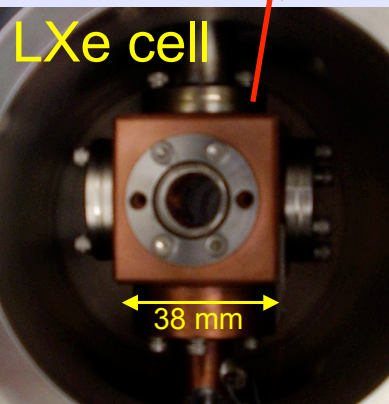
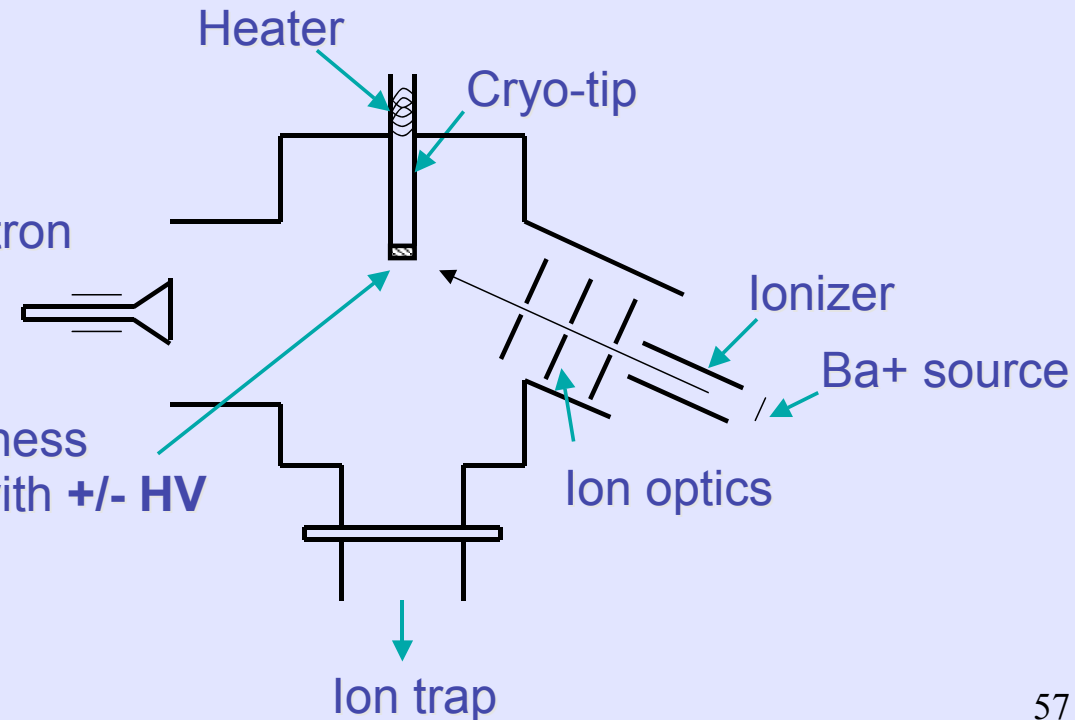
In progress / to come

Tip moving robot



- Freeze Ba^+ from a source on thin Xe ice
- Interface cryo-tip with linear ion trap

Channeltron



LXe cell

38 mm

Ion trap

Ion trap

Main challenges

- Understanding efficiency of Ba^+ transfer from LXe bath into ion trap
- Understand ionic state of Ba in LXe (with ionization cloud following 2β decay), on tip, and after release

Towards a ton-scale EXO detector

- grabbing setup in large LXe bath (spatial resolution)
- need low trigger rate ($\sim 1/\text{hour}$): ultra low background is still mandatory
- use 2v channel for calibration
- energy resolution becomes the limiting factor: work as hard as possible to improve it
- consider all Ba-producing sources, if any

EXO projected sensitivity

Assumptions:

- 1) 80% enrichment in 136
- 2) Intrinsic low background + Ba tagging eliminate all radioactive background
- 3) Energy res only used to separate the 0 ν from 2 ν modes:
Select 0 ν events in a $\pm 2\sigma$ interval centered around the 2.481MeV endpoint
- 4) Use for 2 $\nu\beta\beta$ $T_{1/2} > 1 \cdot 10^{22}$ yr (Bernabei et al. measurement)

Case	Mass (ton)	Eff. (%)	Run Time (yr)	σ_E/E @ 2.5MeV (%)	2 $\nu\beta\beta$ Background (events)	$T_{1/2}^{0\nu}$ (yr, 90% CL)	Majorana mass (meV) QRPA \ddagger (NSM) $\#$	
Conservative	1	70	5	1.6*	0.5 (use 1)	$2 \cdot 10^{27}$	33	(95)
Aggressive	10	70	10	1 \dagger	0.7 (use 1)	$4.1 \cdot 10^{28}$	7.3	(21)

* $\sigma(E)/E = 1.6\%$ obtained in EXO R&D, Conti et al Phys Rev B 68 (2003) 054201

\dagger $\sigma(E)/E = 1.0\%$ considered as an aggressive but realistic guess with large light collection area

\ddagger QRPA: A.Staudt et al. Europhys. Lett.13 (1990) 31; Phys. Lett. B268 (1991) 312

$\#$ NSM: E.Caurier et al. Phys Rev Lett 77 (1996) 1954

Conclusions

- the $^{136}\text{Xe}/^{136}\text{Ba}$ system is very attractive for measuring neutrinoless double beta decay virtually without background other than the 2ν channel
- EXO-200 is well under way: it will likely measure the lifetime of the $2\nu\beta\beta$ decay of ^{136}Xe , measure (if neutrino masses are degenerate) or set competitive limit for $0\nu\beta\beta$ of ^{136}Xe , and provide invaluable input for a larger EXO detector (backgrounds, technical solutions, materials, Xe purity in large detectors)
- Ba^+ identification scheme is well defined, and a grabbing sequence is currently being developed
- A ton-scale EXO detector could be realized in a not-so-far future



Collaboration

D. Leonard, A. Piepke
Physics Department, University of Alabama, Tuscaloosa AL, USA

P. Vogel
Physics Department, Caltech, Pasadena CA, USA

C. Hargrove, D. Sinclair, V. Strickland
Physics Department, Carleton University, Ottawa QC, Canada

W. Fairbank Jr., K. Hall, B. Mong
Physics Department, Colorado State University, Fort Collins CO, USA

M.Moe
Physics Department, UC Irvine, Irvine CA, USA

D. Akimov, A. Burenkov, M. Danilov, A. Dolgolenko, A. Kovalenko, D. Kovalenko, G. Smirnov, V. Stekhanov
ITEP Moscow, Russia

J. Farine, D. Hallman, C. Virtue
Physics Department, Laurentian University, Sudbury ON, Canada

E. Baussan, M. Hauger, F. Juget, Y. Martin, L. Ounalli, D. Schenker, J-L.Vuilleumier, J-M.Vuilleumier, P.Weber
Physics Department, University of Neuchatel, Switzerland

C.Hall, L.Kaufman
Physics Department, University of Maryland, College Park MD, USA

M. Breidenbach, R. Conley, R. Herbst, J. Hodgson, D. Mackay, A. Odian, C. Prescott, P. Rowson, K. Skarpaas, K. Wamba
SLAC, Menlo Park CA, USA

R. DeVoe, P. Fierlinger, B. Flatt, G. Gratta, M. Green, F. LePort, M. Montero Diez, R. Neilson, K. O'Sullivan, A. Pocar, J. Wodin
Physics Department, Stanford University, Stanford CA, USA

